# Periodic Solutions to a Cahn-Hilliard-Willmore Equation in the Plane 

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#### Abstract

In this paper we construct entire solutions to the phase field equation of Willmore type $-\Delta\left(-\Delta u+W^{\prime}(u)\right)+W^{\prime \prime}(u)\left(-\Delta u+W^{\prime}(u)\right)=0$ in the Euclidean plane, where $W(u)$ is the standard double-well potential $\frac{1}{4}\left(1-u^{2}\right)^{2}$. Such solutions have a non-trivial profile that shadows a Willmore planar curve, and converge uniformly to $\pm 1$ as $x_{2} \rightarrow \pm \infty$. These solutions give a counterexample to the counterpart of Gibbons' conjecture for the fourth-order counterpart of the Allen-Cahn equation. We also study the $x_{2}$-derivative of these solutions using the special structure of Willmore's equation.

\section*{Contents} 1. Introduction . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . 822 2. Planar Willmore Curves . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . 825 2.1. Existence of Willmore Curves . . . . . . . . . . . . . . . . . . . . . . . . 826 2.2. The Linearized Problem . . . . . . . . . . . . . . . . . . . . . . . . . . . . 827 3. Approximate Solutions . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . 831 3.1. Fermi Coordinates Near $\gamma_{T}$. . . . . . . . . . . . . . . . . . . . . . . . . . 831 3.2. The Laplacian in Fermi Coordinates . . . . . . . . . . . . . . . . . . . . . 832 3.3. Construction of the Approximate Solution . . . . . . . . . . . . . . . . . . 834 4. The Lyapunov-Schmidt Reduction . . . . . . . . . . . . . . . . . . . . . . . . 841 4.1. The Auxiliary Equation: A Gluing Procedure . . . . . . . . . . . . . . . . 841 4.2. The Bifurcation Equation . . . . . . . . . . . . . . . . . . . . . . . . . . . 844 4.3. Proof of Theorem 1 . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . 849 5. Proof of Some Technical Results . . . . . . . . . . . . . . . . . . . . . . . . . 851 5.1. Proof of Proposition 9 . . . . . . . . . . . . . . . . . . . . . . . . . . . . . 851 5.1.1 The Linear Problem . . . . . . . . . . . . . . . . . . . . . . . . . . 851 5.1.2 A Fixed Point Argument . . . . . . . . . . . . . . . . . . . . . . . . 855 5.2. Proof of Proposition 10 . . . . . . . . . . . . . . . . . . . . . . . . . . . . 855 5.2.1 The Linear Problem . . . . . . . . . . . . . . . . . . . . . . . . . . 856 5.2.2 Proof of Proposition 10 Completed . . . . . . . . . . . . . . . . . . 860 6. Appendix . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . 861

References . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . . 864


## 1. Introduction

The Cahn-Hilliard equation was introduced in [14] to model phase separation of binary fluids. Typically, in experiments, a mixture of fluids tends to gradually self-arrange into more regular oscillatory patterns, with a sharp transition from one component to the other. Applications of this model include complex fluids and soft matter, such as are found in polymer science. The goal of this paper is to rigorously construct planar solutions modelling wiggly transient patterns exhibited by a Cahn-Hilliard-Willmore equation, and to relate them to some existing literature concerning the Allen-Cahn equation, a second-order counterpart describing phase separation in alloys.

Let us begin by recalling some basic features about the Allen-Cahn equation

$$
\begin{equation*}
-\Delta u=u-u^{3} \tag{1}
\end{equation*}
$$

introduced in [5]. Here $u$ represents, up to an affine transformation, the density of one of the components of an alloy, whose energy per unit volume is given by a double-well potential $W$

$$
\begin{equation*}
W(u)=\frac{1}{4}\left(1-u^{2}\right)^{2} . \tag{2}
\end{equation*}
$$

Global minimizers (for example taken among functions with a prescribed average) of the integral of $W$ consist of the functions attaining only the values $\pm 1$. Since of course this set of functions has no structure whatsoever, usually a regularization of the energy of the following type is considered:

$$
E_{\varepsilon}(u)=\int_{\Omega}\left(\frac{\varepsilon}{2}|\nabla u|^{2}+\frac{\left(1-u^{2}\right)^{2}}{4 \varepsilon}\right) \mathrm{d} x
$$

which penalizes too frequent phase transitions.
It was shown in [32] that under suitable assumptions $E_{\varepsilon}$ Gamma-converges as $\varepsilon \rightarrow 0$ to the perimeter functional and therefore its critical points are expected to have transitions approximating surfaces with zero mean curvature. In particular, minimizers for $E_{\varepsilon}$ should produce interfaces that are stable minimal surfaces, see [31,38] (and also [26]). The relation between the stability of solutions to (1) and their monotonicity has been the subject of several investigations, see for example [3, 24,37]. In particular, a celebrated conjecture by DE GIORGI [20] states that solutions to (1) that are monotone in some direction should depend on one variable only in dimension $n \leqq 8$. This restriction on $n$ is crucial, since in large dimension there exist stable minimal surfaces that are not planar, see [12], and recently some entire solutions modelled on them were constructed in [18]. Further solutions with nontrivial profiles were produced for example in [2,13,16,17,19].

Another related conjecture named as the Gibbons conjecture, motivated by problems in cosmology, asserts that solutions to (1) such that

$$
u\left(x^{\prime}, x_{n}\right) \rightarrow \pm 1 \quad \text { as } x_{n} \rightarrow \pm \infty \quad \text { uniformly for } x^{\prime} \in \mathbb{R}^{n-1}
$$

should also be one-dimensional. This conjecture was indeed fully proved in all dimensions, see [7,11,22,24].

We turn next to the geometric equation

$$
\begin{equation*}
-\Delta\left(-\Delta u+W^{\prime}(u)\right)+W^{\prime \prime}(u)\left(-\Delta u+W^{\prime}(u)\right)=0 \tag{3}
\end{equation*}
$$

which from now on we will refer to as the Cahn-Hilliard-Willmore equation.
Similarly to (1), also this equation is variational; introducing a scaling parameter $\varepsilon>0$, its Euler-Lagrange functional is given by

$$
\mathcal{W}_{\varepsilon}(u)=\frac{1}{2 \varepsilon} \int_{\Omega}\left(\varepsilon \Delta u-\frac{W^{\prime}(u)}{\varepsilon}\right)^{2} \mathrm{~d} x
$$

Notice that when the integrand vanishes identically $u$ solves a scaled version of (1). As for $E_{\varepsilon}, \mathcal{W}_{\varepsilon}$ also has a geometric interpretation as $\varepsilon \rightarrow 0$. Although the characterization of Gamma-limit is not as complete as for the Allen-Cahn equation, some partial results are known about convergence to (a multiple of) the Willmore energy of the limit interface, that is the integral of the mean curvature squared

$$
\mathcal{W}_{0}(u)=\int_{\partial E \cap \Omega} H_{\partial E}^{2}(y) \mathrm{d} \mathcal{H}^{N-1}
$$

In [10], Bellettini and Paolini proved the $\Gamma$ - lim sup inequality for smooth Willmore hypersurfaces, while the $\Gamma$ - lim inf inequality has been proved in dimension $N=2,3$ by Röger and Schätzle in [36], and, independently, in dimension $N=2$, by Nagase and Tonegawa in [33]. It is an open problem to study in a higher dimension, as well as to understand for which class of sets the Gamma-limit might exist.

Apart from the relation to the Cahn-Hilliard-Willmore energy, the Willmore functional appears as bending energy of plates and membranes in mechanics and in biology, and it also enters in general relativity as the Hawking mass of a portion of space-time. This energy also has interest in geometry, since it is invariant under Möbius transformations. Critical surfaces of $\mathcal{W}$ are called Willmore hypersurfaces, and they are known to exist for any genus, see [8]. The Euler equation satisfied by this kind of hypersurface is

$$
-\Delta_{\Sigma} H=\frac{1}{2} H^{3}-2 H K
$$

Interesting Willmore surfaces are Clifford tori (and their Möbius transformations), which can be obtained by rotating around the $z$-axis a circle of radius 1 and with center at distance $\sqrt{2}$ from the axis. Due to a recent result in [30], establishing the so-called Willmore conjecture, this torus minimizes the Willmore energy among all surfaces of positive genus. In [35], up to a small Lagrange multiplier, solutions of (3) in $\mathbb{R}^{3}$ were found with interfaces approaching a Clifford torus, converging to -1 in its interior and to +1 on its exterior.

Here we will show the existence of solutions to (3) in the plane with an interface periodic in $x_{1}$, shadowing a $T$-periodic (in the arc-length parameter) Willmore curve $\gamma_{T}$, whose profile is given in Fig. 1. Notice that, by the vanishing of Gaussian


Fig. 1. The Willmore curve $\gamma_{T}$
curvature of cylindrical surfaces, for one-dimensional curves the Willmore equation reduces to an ODE for the planar curvature, namely

$$
k^{\prime \prime}=-\frac{1}{2} k^{3}
$$

This equation can be explicitly solved using special functions, and then integrated to produce the above Willmore curves $\gamma_{T}$. Indeed, every non-affine complete planar Willmore curve coincides, up to an affine transformation with the curve $\gamma_{T}$, see [29, p. 3907].

Apart from producing a first non-compact profile of this type for the equation, our aim is to explore the relation between one-dimensionality of solutions and their limit properties. In fact, our construction shows that the straightforward counterpart of Gibbons' conjecture for (1) is false. Our main result reads as follows:

Theorem 1. There exists $T_{0}>0$ such that, for any $T>T_{0}$ there exists a $T$-periodic planar Willmore curve $\gamma_{T}$ and a solution $u_{T}$ to

$$
-\Delta\left(-\Delta u_{T}+W^{\prime}\left(u_{T}\right)\right)+W^{\prime \prime}\left(u_{T}\right)\left(-\Delta u_{T}+W^{\prime}\left(u_{T}\right)\right)=0
$$

such that

$$
u_{T}\left(x_{1}, x_{2}\right) \rightarrow \pm 1 \quad \text { as } x_{2} \rightarrow \pm \infty \quad \text { uniformly for } x_{1} \in \mathbb{R}
$$

and

$$
\begin{equation*}
\operatorname{dist}\left(\gamma_{T},\left\{x \in \mathbb{R}^{2}: u_{T}(x)=0\right\}\right)<\frac{c}{T} \tag{4}
\end{equation*}
$$

The function $u_{T}$ also satisfies the symmetries

$$
u_{T}\left(x_{1}, x_{2}\right)=-u_{T}\left(-x_{1},-x_{2}\right)=-u_{T}\left(x_{1}+\frac{L}{2},-x_{2}\right) \quad \text { for every } x \in \mathbb{R}^{2}
$$

In particular, it is L-periodic in the $x_{1}$ variable, where $L:=\left(\gamma_{T}\right)_{1}(T)-\left(\gamma_{T}\right)_{1}(0)>$ 0 , and furthermore, there exists a fixed constant $C_{0}$ such that

$$
\begin{equation*}
\partial_{x_{2}} u_{T}\left(x_{1}, x_{2}\right) \geqq-C_{0} T^{-3} \quad \text { for all }\left(x_{1}, x_{2}\right) \in \mathbb{R}^{2} \text { and all } T>T_{0} \tag{5}
\end{equation*}
$$

In the literature there are nowadays several constructions of interfaces starting from given limit profiles via Lyapunov-Schmidt reductions, see the abovementioned references. However, being the Cahn-Hilliard-Willmore equation of fourth order, here one needs a rather careful expansion using also smoothing operators. Moreover, to the best of our knowledge, our solution seems to be the first one in the literature with a non-compact (and non-trivial) transition profile for (3).

Notice also that the curve $\gamma_{T}$ is vertical at an equally-spaced sequence of points lying on the $x$-axis. Therefore the gradient of $u_{T}$ is nearly horizontal at these points, and it is quite difficult to understand the monotonicity (in $x_{2}$ ) of the solutions in these regions. Apart from the fact that the equation is of fourth-order, and hence rather involved to analyse, we need to expand a formally (3) up to the fifth order in $\frac{1}{T}$ for proving the estimate (5). In practice, we need to find a sufficiently good approximate solution to (3) by adding suitable corrections to a naive transition layer along $\gamma_{T}$, and then by tilting properly the transition profile by a $T$-periodic function $\phi$. This tilting, which is of order $O\left(T^{-1}\right)$, satisfies a linearized Willmore equation of the form

$$
\tilde{L}_{0} \bar{\phi}=F,
$$

where $F(t)$ is an explicit function of the curvature of $\gamma_{T}$ and its derivatives. The special structure of the right-hand side in our case and the special structure of the Willmore equation make it possible to find an explicit solution (again, in terms of special functions) for $\bar{\phi}$, depending only on the curvature of $\gamma_{T}$ and its derivatives.

Remark 2. Unfortunately the main order term $\bar{\phi}$ in the perpendicular tilting of the interface with respect to $\gamma_{T}$ is flat at its vertical points, so we can neither claim a full monotonicity of the solutions, nor disprove it. With our analysis and some extra work it should be possible to prove monotonicity of $u_{T}$ in suitable portions of the plane, however to understand the monotonicity near those special points one would need either much more involved expansions and/or different ideas.

The plan of the paper is as follows. In Section 2 we study planar Willmore curves, and analyse some properties, including the spectral ones, of the linearised Willmore equation. In Section 3 we construct approximate solutions, expanding (3) up to the fifth order in $T^{-1}$, in order to understand the normal tilting of the interface to $\gamma_{T}$. In Section 4 we give the outline of the proof of our main result, performing a Lyapunov-Schmidt reduction of the problem on the normal tilting $\phi$. Sections 5 and the "Appendix" are devoted to the proofs of some technical results; the former, concerning the reduction technique, while the latter deals with the main order term $\bar{\phi}$ in the expansion of $\phi$.

## 2. Planar Willmore Curves

In this section we collect some material about the existence of planar Willmore curves, analysing then their spectral properties with respect to the second variation of the Willmore energy.

Recall that the Willmore energy of a curve $\gamma:[0,1] \rightarrow \mathbb{R}^{2}$ is defined as the integral of the curvature squared

$$
\int_{\gamma} k(s)^{2} \mathrm{~d} s
$$

Extremizing with respect to variations that are compactly supported in $(0,1)$, one finds that critical points satisfy the Willmore equation

$$
\begin{equation*}
k^{\prime \prime}=-\frac{1}{2} k^{3} \tag{6}
\end{equation*}
$$

### 2.1. Existence of Willmore Curves

Recall first the definition of the Jacobi cosine function, see for example [9]. For $m \in(0,1)$ define

$$
\sigma(\varphi, m)=\int_{0}^{\varphi} \frac{\mathrm{d} \theta}{\sqrt{1-m \sin ^{2} \theta}}
$$

and then implicitly define the function by

$$
\begin{equation*}
\mathrm{cn}(\sigma(\varphi, m) \mid m)=\cos \varphi \tag{7}
\end{equation*}
$$

The above equation (6) admits (only) periodic solutions that, up to a dilation and translation, are given by

$$
\begin{equation*}
k(s)=\sqrt{2} \mathrm{cn}\left(s+\bar{T} / 4 \left\lvert\, \frac{1}{2}\right.\right) \tag{8}
\end{equation*}
$$

For this choice, the period $\bar{T}$ has the approximate value $\bar{T} \simeq 10.488$. Using the conservation of Hamiltonian energy, this function satisfies

$$
\begin{equation*}
\left(k^{\prime}\right)^{2}(s)=-\frac{k^{4}(s)}{4}+1 \tag{9}
\end{equation*}
$$

The above function can be integrated to produce a Willmore curve, by the formula (with an abuse of notation, we will always use the same letter both for the curve and for its parametrization)

$$
\gamma(s)=\int_{0}^{s}\binom{-\sin \left(\int_{0}^{s} k(\tau) \mathrm{d} \tau\right)}{\cos \left(\int_{0}^{s} k(\tau) \mathrm{d} \tau\right)} \mathrm{d} s
$$

Notice that $\gamma$ is parametrized by arc length. If we set $\gamma_{T}(s):=\varepsilon^{-1} \gamma(\varepsilon s)$, for $\varepsilon=\bar{T} / T$, then it is still true that $\left|\gamma_{T}^{\prime}(s)\right|=1$ for any $s \in \mathbb{R}$. In other words, $\gamma_{T}$ also denotes the rescaled curve $\left\{\varepsilon^{-1} \zeta: \zeta \in \gamma\right\}$, still parametrized by arc length. Our aim is to construct solutions $u_{T}$ with a transition layer close to $\gamma_{T}$, that are odd and periodic in $x_{1}$ and fulfilling the symmetry property

$$
\begin{align*}
u_{T}\left(x_{1}, x_{2}\right) & =-u_{T}\left(-x_{1},-x_{2}\right)=-u_{T}\left(x_{1}+\frac{L}{2},-x_{2}\right), \\
L & :=\left(\gamma_{T}\right)_{1}(T)-\left(\gamma_{T}\right)_{1}(0) . \tag{10}
\end{align*}
$$

The curvature of $\gamma_{T}$ is defined by

$$
k_{\varepsilon}(s):=-\left\langle\gamma_{T}^{\prime}(s), \gamma_{T}^{\prime \prime}(s)^{\perp}\right\rangle, \quad w^{\perp}=\left(-w_{2}, w_{1}\right)
$$

and clearly by the arc-length parametrisation one has $\gamma_{T}^{\prime \prime}(s)=k_{\varepsilon}(s) \gamma_{T}^{\prime}(s)^{\perp}$. In what follows, when the subscript $\varepsilon$ is omitted, it will be assumed to be equal to 1 , that is we will set $\gamma:=\gamma_{1}, k:=k_{1}$, etc.

### 2.2. The Linearized Problem

We discuss next the linearization of the Willmore equation, namely we consider the problem

$$
\begin{equation*}
\tilde{L}_{0} \phi=g \text { in } \mathbb{R}, \tag{11}
\end{equation*}
$$

where $g: \mathbb{R} \rightarrow \mathbb{R}$ is a given $\bar{T}$-periodic function. Recall from formula (33) in [27] that $\tilde{L}_{0}$ is given by
$\tilde{L}_{0} \phi=\phi^{(4)}+\left(\frac{5}{2} k^{2} \phi^{\prime}\right)^{\prime}+\left(3\left(k^{\prime}\right)^{2}-\frac{1}{2} k^{4}\right) \phi=\phi^{(4)}+\left(\frac{5}{2} k^{2} \phi^{\prime}\right)^{\prime}+\left(3-\frac{5}{4} k^{4}\right) \phi$,
where the conservation law (9) has been used. Given the symmetries of the problem, we are interested in right-hand sides $g$ that satisfy the following conditions:

$$
g(s)=-g(-s)=-g(s+\bar{T} / 2)
$$

hence we define the spaces

$$
\begin{equation*}
C_{\bar{T}}^{n, \alpha}(\mathbb{R}):=\left\{\phi \in C^{n, \alpha}(\mathbb{R}): \phi(s)=-\phi(-s)=-\phi\left(s+\frac{\bar{T}}{2}\right)\right\} \tag{13}
\end{equation*}
$$

where $\bar{T}>0, n \geqq 0$ is an integer, $0<\alpha<1$ and $C^{n, \alpha}(\mathbb{R})$ is the space of functions $\phi: \mathbb{R} \rightarrow \mathbb{R}$ that are $n$ times differentiable and whose $n$th derivative is Hölder continuous of exponent $\alpha$. We endow the spaces $C_{\bar{T}}^{n, \alpha}(\mathbb{R})$ with the norms

$$
\|\phi\|_{C^{n, \alpha}(\mathbb{R})}=\sum_{j=0}^{n}\left\|\nabla^{j} \phi\right\|_{L^{\infty}(\mathbb{R})}+\sup _{s \neq t} \frac{\left|\phi^{(n)}(s)-\phi^{(n)}(t)\right|}{|t-s|^{\alpha}}
$$

Roughly speaking, these spaces consist of functions that respect the symmetries of the curve $\gamma$, in the sense that they are even, periodic with period $\bar{T}$, and they change sign after a translation of half a period. We have then the following result:

Proposition 3. Let $\bar{T}>0$. Let $g \in C^{0, \alpha}(\mathbb{R})$ satisfy $g(s)=-g(-s)=-g(s+$ $\bar{T} / 2)$ for all $s \in \mathbb{R}$. Then there is a unique function $\phi \in C^{4, \alpha}(\mathbb{R})$ that solves equation (11) and satisfies

$$
\phi(s)=-\phi(-s)=-\phi(s+\bar{T} / 2) \quad \forall s \in \mathbb{R}
$$

Moreover, the estimate $\|\phi\|_{C^{4, \alpha}(\mathbb{R})} \leqq c\|g\|_{C^{0, \alpha}(\mathbb{R})}$ holds for some positive number $c$ independent of $g$.

Proof. By considering extensions by periodicity, it is sufficient to prove unique resolvability of $\tilde{L}_{0} \phi=g$ on $[0, \bar{T}]$ for $\phi$ in the space

$$
\begin{aligned}
C_{\bar{T}, 0}^{4, \alpha}([0, \bar{T}]) & :=\left\{\left.\phi\right|_{[0, \bar{T}]}: \phi \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})\right\}=\left\{\phi \in C^{4, \alpha}([0, \bar{T}]): \phi(s)\right. \\
& =-\phi(\bar{T}-s)=-\phi(s+\bar{T} / 2) \text { for } 0 \leqq s \leqq \bar{T} / 2\}
\end{aligned}
$$

We observe that, by construction, any function $\phi \in C_{\bar{T}, 0}^{4, \alpha}([0, \bar{T}])$ satisfies $\phi^{(j)}(0)=$ $\phi^{(j)}(\bar{T}), 0 \leqq j \leqq 4$, hence we can extend $\phi$ to a function in $C_{\bar{T}}^{4, \alpha}(\mathbb{R})$.

We denote by $(-\Delta)^{-2} h$ the unique solution $\phi$ to

$$
\begin{cases}\phi^{(4)}=h & \text { on }[0, \bar{T}] \\ \phi(0)=\phi(\bar{T})=\phi^{\prime \prime}(0)=\phi^{\prime \prime}(\bar{T})=0 & \text { (homogeneous Navier boundary conditions) }\end{cases}
$$

where $h \in C_{\bar{T}}^{0, \alpha}(\mathbb{R})$ is given. Such a solution $\phi$ is in $C^{4, \alpha}([0, \bar{T}])$ and fulfils the estimate

$$
\|\phi\|_{C^{4, \alpha}(\mathbb{R})}=\|\phi\|_{C^{4, \alpha}(0, \bar{T})} \leqq c\|h\|_{C^{0, \alpha}(0, \bar{T})}=c\|h\|_{C^{0, \alpha}(\mathbb{R})}
$$

Since $h$ verifies the symmetries $h(s)=-h(\bar{T}-s)=-h(s+\bar{T} / 2)$, for any $0 \leqq s \leqq \bar{T} / 2$, then also does $\phi$, thus $\phi \in C_{\bar{T}, 0}^{4, \alpha}([0, \bar{T}])$.

Then $\tilde{L}_{0} \phi=g$ is equivalent to

$$
\phi+(-\Delta)^{-2}\left(\left(\frac{5}{2} k^{2} \phi^{\prime}\right)^{\prime}+\left(3-\frac{5}{4} k^{4}\right) \phi\right)=(-\Delta)^{-2}(g)
$$

In particular, the Fredholm alternative in the space $C_{\bar{T}, 0}^{4, \alpha}([0, \bar{T}])$ applies, so the equation is solvable for every $g \in C_{\bar{T}}^{0, \alpha}(\mathbb{R})$ if and only if the homogeneous problem is uniquely solvable.

Exploiting the symmetries of the Willmore equation, if $\nu_{\gamma}$ denotes the normal vector to the curve $\gamma$, the following four functions represent Jacobi fields for the linearized Willmore equation:

$$
\begin{aligned}
& \psi_{1}=\left\langle(0,-1), v_{\gamma}\right\rangle ; \quad \psi_{2}=\left\langle(1,0), v_{\gamma}\right\rangle ; \quad \psi_{3}=\left\langle\left(\gamma_{2},-\gamma_{1}\right), v_{\gamma}\right\rangle \\
& \psi_{4}=\left\langle\left(\gamma_{1}, \gamma_{2}\right), v_{\gamma}\right\rangle
\end{aligned}
$$

Using the fact that $\nu_{\gamma}=\left(\gamma_{2}^{\prime},-\gamma_{1}^{\prime}\right)$ (here $\gamma_{1}$ and $\gamma_{2}$ represent the horizontal and vertical components of $\gamma$ ) one finds that

$$
\left(\psi_{1}, \ldots, \psi_{4}\right):=\left(\gamma_{1}^{\prime}, \gamma_{2}^{\prime}, \gamma_{1} \gamma_{1}^{\prime}+\gamma_{2} \gamma_{2}^{\prime}, \gamma_{1} \gamma_{2}^{\prime}-\gamma_{2} \gamma_{1}^{\prime}\right)
$$

We claim that these functions are linearly independent: indeed, using $k^{\prime}(0)=$ $k^{\prime}(\bar{T})=-1$ and $\gamma_{1}(\bar{T}) \neq 0$ we get

$$
\begin{aligned}
& \left(\begin{array}{c}
\psi_{1}(0) \\
\psi_{1}(\bar{T}) \\
\psi_{1}^{\prime \prime}(0) \\
\psi_{1}^{\prime \prime}(\bar{T})
\end{array}\right)=\left(\begin{array}{l}
0 \\
0 \\
1 \\
1
\end{array}\right), \quad\left(\begin{array}{l}
\psi_{2}(0) \\
\psi_{2}(\bar{T}) \\
\psi_{2}^{\prime \prime}(0) \\
\psi_{2}^{\prime \prime}(\bar{T})
\end{array}\right)=\left(\begin{array}{l}
1 \\
1 \\
0 \\
0
\end{array}\right) \\
& \left(\begin{array}{c}
\psi_{3}(0) \\
\psi_{3}(\bar{T}) \\
\psi_{3}^{\prime \prime}(0) \\
\psi_{3}^{\prime \prime}(\bar{T})
\end{array}\right)=\left(\begin{array}{c}
0 \\
0 \\
0 \\
\gamma_{1}(\bar{T})
\end{array}\right), \quad\left(\begin{array}{c}
\psi_{4}(0) \\
\psi_{4}(\bar{T}) \\
\psi_{4}^{\prime \prime}(0) \\
\psi_{4}^{\prime \prime}(\bar{T})
\end{array}\right)=\left(\begin{array}{c}
0 \\
\gamma_{1}(\bar{T}) \\
0 \\
0
\end{array}\right)
\end{aligned}
$$

Being the homogeneous ODE $\tilde{L}_{0} \phi=0$ of fourth order in $\phi$, all its $\bar{T}$-periodic solutions are spanned by $\left\{\psi_{1}, \psi_{2}, \psi_{3}, \psi_{4}\right\}$. From the above formulas one infers that a function $\phi \in C_{\bar{T}, 0}^{4, \alpha}([0, \bar{T}])$ that is a linear combination $\left(\psi_{1}, \psi_{2}, \psi_{3}, \psi_{4}\right)$ satisfies homogeneous Navier boundary conditions if and only if it is trivial. Hence the homogeneous problem has only the trivial solution, and the equation $\tilde{L}_{0} \phi=g$ (with the desired boundary conditions) is uniquely solvable in $C_{\bar{T}, 0}^{4, \alpha}([0, \bar{T}])$, as claimed. The norm estimate follows from

$$
\|\phi\|_{C^{4, \alpha}(\mathbb{R})}=\|\phi\|_{C^{4, \alpha}(0, \bar{T})} \leqq c\|g\|_{C^{0, \alpha}(0, \bar{T})}=c\|g\|_{C^{0, \alpha}(\mathbb{R})},
$$

where the inequality results from higher order Schauder estimates, see example [23]. Notice that reducing the problem on $\mathbb{R}$ to a problem on $[0, \bar{T}]$ ensures compactness.

We need next to invert the linearized operator for a specific right-hand side, arising from high-order expansion (in $\varepsilon=\frac{\bar{T}}{T}$ ) of the approximate solutions, see Section 3. We have the following result:

Proposition 4. Let

$$
\bar{g}(s)=\frac{9}{8} k(s)^{5}-9 k(s) k^{\prime}(s)^{2}
$$

Then the equation $\tilde{L}_{0} \bar{\phi}=\bar{g}$ admits a unique smooth solution $\bar{\phi} \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$ which additionally satisfies
(i) $\bar{\phi}(s) k(s) \geqq 0$ for all $s \in \mathbb{R}$,
(ii) $\bar{\phi}^{\prime}(s)=0$ whenever $k(s)=0$.

Remark 5. The solution $\bar{\phi}$ can be written explicitly in terms of hyper-geometric functions, see Chapter 15 in [1] or [6] for the notation we are using and additional properties. Indeed, for every $\mu_{0}, \mu_{1} \in \mathbb{R}$, a formal solution $\bar{\phi}$ is given by the formula $\bar{\phi}(s)=\Phi(k(s))$, where

$$
\Phi(z):=\mu_{0} z+\mu_{1} z^{3}+\frac{37-40 \mu_{0}}{960} z^{5}+r_{1}(z)+r_{2}(z)+r_{3}(z)
$$

for functions $r_{1}, r_{2}, r_{3}$ given by

$$
\begin{aligned}
& r_{1}(z):=\frac{\mu_{1}}{28} z^{7}{ }_{2} F_{1}\left(1, \frac{5}{4} ; \frac{11}{4} ; \frac{z^{4}}{4}\right), \quad r_{2}(z):=-\frac{41}{640} z^{9}{ }_{2} F_{1}\left(1, \frac{7}{4} ; \frac{13}{4} ; \frac{z^{4}}{4}\right), \\
& r_{3}(z):=\left(-\frac{3}{896} \mu_{0}+\frac{465}{7168}\right) z^{9}{ }_{3} F_{2}\left(1, \frac{7}{4}, \frac{5}{2} ; \frac{11}{4}, 3 ; \frac{z^{4}}{4}\right)
\end{aligned}
$$

This representation, however, does not seem to be helpful when discussing the regularity properties of the function $\Phi \circ k$ as we will discuss in the proof below.

Proof. Since we are looking for an odd solution $\bar{\phi} \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$, motivated by the special features of Willmore's equation, we consider the ansatz $\bar{\phi}(s)=\Phi(k(s))$. After some calculations it is possible to write $\tilde{L}_{0} \bar{\phi}=\bar{g}$ as an ODE for $\Phi$, namely

$$
\begin{align*}
& \frac{1}{16}\left(\left(z^{4}-4\right)^{2} \Phi^{(4)}(z)+12 s^{3}\left(z^{4}-4\right) \Phi^{(3)}(z)+2 z^{2}\left(13 z^{4}-28\right) \Phi^{\prime \prime}(z)\right.  \tag{14}\\
& \left.\quad-16\left(z^{4}-2\right) z \Phi^{\prime}(z)+\left(48-20 z^{4}\right) \Phi(z)\right)=-9 z+\frac{27}{8} z^{5}
\end{align*}
$$

This equation can be solved explicitly in terms of a series $\Phi(z)=\sum_{k=0}^{\infty} \mu_{k} z^{2 k+1}$, where the parameters $\mu_{0}, \mu_{1}$ are free and $\mu_{2}, \mu_{3}, \ldots$ are determined recursively by the above ODE. Their precise definitions are provided in the "Appendix", see (85). This series has convergence radius $\sqrt{2}$ so it is not clear a priori whether $s \mapsto \bar{\phi}(s)=\Phi(k(s))$ defines a function of $C_{\bar{T}}^{4, \alpha}(\mathbb{R})$. In order to ensure this we impose that the solution $\bar{\phi}$ is even about $-\bar{T} / 4$ and $\bar{T} / 4$, that is we require that

$$
\begin{equation*}
\bar{\phi}^{\prime}(s) \rightarrow 0, \quad \bar{\phi}^{\prime \prime \prime}(s) \rightarrow 0 \quad \text { as }|s| \rightarrow \bar{T} / 4 \tag{15}
\end{equation*}
$$

Notice that $k(s)$ converges to $\sqrt{2}$ as $|s| \rightarrow \frac{T}{4}$. The calculations from the "Appendix" show that (15) holds if and only if we choose

$$
\begin{equation*}
\mu_{0}=0, \quad \mu_{1}=\frac{\pi^{2}}{8 \Gamma\left(\frac{3}{4}\right)^{4}} \tag{16}
\end{equation*}
$$

The corresponding solution is given by

$$
\bar{\phi}(s)=\frac{3 \pi \sqrt{2}}{64 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \sum_{m=0}^{\infty}\left(-\frac{\Gamma\left(m+\frac{3}{4}\right)}{2 \cdot 4^{m} \Gamma\left(m+\frac{9}{4}\right)} k(s)^{4^{m+5}}+\frac{\Gamma\left(m+\frac{1}{4}\right)}{4^{m} \Gamma\left(m+\frac{7}{4}\right)} k(s)^{4 m+3}\right) .
$$

Thanks to (15) this solution can be reflected evenly about $s= \pm \bar{T} / 4$, so we obtain by standard arguments that $\bar{\phi} \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$. From the above formula we find $\bar{\phi}^{\prime}(s)=O\left(k^{\prime}(s) k(s)^{2}\right) \rightarrow 0$ as $k(s) \rightarrow 0$ as well as

$$
\begin{aligned}
\bar{\phi}(s) k(s) & =\frac{3 \pi \sqrt{2} k(s)^{4}}{64 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \sum_{m=0}^{\infty}\left(-\frac{\Gamma\left(m+\frac{3}{4}\right)}{\Gamma\left(m+\frac{9}{4}\right)} \cdot \frac{k(s)^{2}}{2}+\frac{\Gamma\left(m+\frac{1}{4}\right)}{\Gamma\left(m+\frac{7}{4}\right)}\right)\left(\frac{k(s)^{4}}{4}\right)^{m} \\
& \geqq \frac{3 \pi \sqrt{2} k(s)^{4}}{64 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \sum_{m=0}^{\infty} \underbrace{\left(-\frac{\Gamma\left(m+\frac{3}{4}\right)}{\Gamma\left(m+\frac{9}{4}\right)}+\frac{\Gamma\left(m+\frac{1}{4}\right)}{\Gamma\left(m+\frac{7}{4}\right)}\right)}_{\geqq 0}\left(\frac{k(s)^{4}}{4}\right)^{m} \geqq 0 .
\end{aligned}
$$

Hence, claims (i) and (ii) are proved and we can conclude.

## 3. Approximate Solutions

In this section we introduce an approximate solution of (2), which we need to expand up to the fifth order in $\varepsilon=\bar{T} / T$. For doing this, we use Fermi coordinates around a perturbation of the curve

$$
\gamma_{T}(\cdot)=\frac{1}{\varepsilon} \gamma(\varepsilon \cdot) ; \quad \varepsilon=\frac{\bar{T}}{T}
$$

( $\gamma$ is the Willmore curve constructed in Section 2) and we expand both the Laplace operator and Cahn-Hilliard-Willmore equation. We also need to add suitable corrections to the approximate solution in order to improve its accuracy; these will allow us to study in more detail the transition curve $\left\{u_{T}=0\right\}$ of the solution constructed in Theorem 1.

### 3.1. Fermi Coordinates Near $\gamma_{T}$

As in [35], we want to use Fermi coordinates near a normal perturbation of the dilated periodic curve $\gamma_{T}$. To this end, we fix $\phi \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$ (recall (13)) such that $\|\phi\|_{C^{4, \alpha}(\mathbb{R})}<1$, and for $T$ large (that is for $\varepsilon=\frac{\bar{T}}{T}$ small) we define the planar map

$$
\begin{equation*}
Z_{\varepsilon}(s, t)=\gamma_{T}(s)+(t+\phi(\varepsilon s)) \gamma_{T}^{\prime}(s)^{\perp}, \quad w^{\perp}:=\left(-w_{2}, w_{1}\right) . \tag{17}
\end{equation*}
$$

Using the fact that $\gamma_{T}:=\varepsilon^{-1} \gamma(\varepsilon s)$ and $\left|\gamma_{T}^{\prime}\right|=1$ we find

$$
\begin{aligned}
\operatorname{det}\left(\partial_{s} Z_{\varepsilon}, \partial_{t} Z_{\varepsilon}\right)= & \operatorname{det}\left(\gamma^{\prime}(\varepsilon s)^{\perp}, \gamma^{\prime}(\varepsilon s)+\varepsilon \phi^{\prime}(\varepsilon s) \gamma^{\prime}(\varepsilon s)^{\perp}\right. \\
& \left. \pm(t+\phi(\varepsilon s)) \varepsilon k(\varepsilon s) \gamma^{\prime}(\varepsilon s)\right) \\
= & \operatorname{det}\left(\gamma^{\prime}(\varepsilon s)^{\perp}, \gamma^{\prime}(\varepsilon s)\right) \cdot(1+\varepsilon k(\varepsilon s)(t+\phi(\varepsilon s))) \\
\geqq & 1-\sqrt{2}(|t|+1) \varepsilon \\
& \geqq \frac{1}{4} \text { for }|t|<\frac{1}{2 \sqrt{2} \varepsilon} .
\end{aligned}
$$

This shows that in the above region the map $Z_{\varepsilon}$ is invertible. Moreover, define

$$
V_{\varepsilon, \phi}:=\left\{x \in \mathbb{R}^{2}: \operatorname{dist}\left(x, \gamma_{T, \phi}\right)<\frac{1}{4 \varepsilon}\right\}
$$

and $\gamma_{T, \phi}:=\gamma_{T}(s)+\phi(\varepsilon s) \gamma_{T}^{\prime}(s)^{\perp}$. Since $\phi \in C^{4, \alpha}(\mathbb{R})$ it follows also that $Z_{\varepsilon}: \mathbb{R} \times$ $\left(-\frac{1}{4 \varepsilon}, \frac{1}{4 \varepsilon}\right) \rightarrow V_{\varepsilon, \phi}$ is a $C^{4, \alpha}$-diffeomorphism. With an abuse of notation, we will write $u$ for $u(s, t)$. We also set

$$
\begin{equation*}
V_{\varepsilon}:=\left\{x \in \mathbb{R}^{2}: \operatorname{dist}\left(x, \gamma_{T}\right)<\frac{1}{4 \varepsilon}\right\} . \tag{18}
\end{equation*}
$$

### 3.2. The Laplacian in Fermi Coordinates

We are interested in the expression of the Laplacian in the above coordinates $(s, t)$. First we assume that $\phi=0$, that is we consider the diffeomorphism

$$
\tilde{Z}_{\varepsilon}: \mathbb{R} \times(-1 / 4 \varepsilon, 1 / 4 \varepsilon) \rightarrow V_{\varepsilon}
$$

defined by

$$
\begin{equation*}
\tilde{Z}_{\varepsilon}(s, z):=\gamma_{T}(s)+z \gamma_{T}^{\prime}(s)^{\perp} \tag{19}
\end{equation*}
$$

The euclidean metric in these coordinates is

$$
g=\left[\begin{array}{cc}
(1-\varepsilon z k(\varepsilon s))^{2} & 0 \\
0 & 1
\end{array}\right]
$$

with determinant $\operatorname{det} g=g_{s s}=(1-\varepsilon z k(\varepsilon s))^{2}$ and inverse given by

$$
g^{-1}=\left[\begin{array}{cc}
(1-\varepsilon z k(\varepsilon s))^{-2} & 0 \\
0 & 1
\end{array}\right]
$$

From now on, the curvature $k$ and its derivatives will always be evaluated at $\varepsilon s$. Using that $g^{s s}=g_{s s}^{-1}$, the Laplacian with respect to this metric is given by

$$
\begin{align*}
\Delta u & =\frac{1}{\sqrt{g_{s s}}} \partial_{s}\left(\sqrt{g_{s s}} g^{s s} \partial_{s} u\right)+\frac{1}{\sqrt{g_{s s}}} \partial_{z}\left(\sqrt{g_{s s}} \partial_{z} u\right) \\
& =\frac{1}{\sqrt{g_{s s}}} \partial_{s}\left(\frac{1}{\sqrt{g_{s s}}} \partial_{s} u\right)+\partial_{z}^{2} u+\frac{\partial_{z} g_{s s}}{2 g_{s s}} \partial_{z} u \\
& =\frac{\partial_{s}^{2} u}{g_{s s}}-\frac{\partial_{s} g_{s s}}{2 g_{s s}^{2}} \partial_{s} u+\partial_{z}^{2} u+\frac{\partial_{z} g_{s s}}{2 g_{s s}} \partial_{z} u . \tag{20}
\end{align*}
$$

Now we compute

$$
\frac{\partial_{z} g_{s s}}{2 g_{s s}}=-\varepsilon \frac{k}{1-\varepsilon z k}
$$

Taylor expanding in $\varepsilon$, we get

$$
\frac{k}{1-\varepsilon z k}=k+\varepsilon k^{2} z+\varepsilon^{2} k^{3} z^{2}+\varepsilon^{3} k^{4} z^{3}+\varepsilon^{4} k^{5} z^{4}+\varepsilon^{5} \bar{h}(\varepsilon s, z)
$$

where the remainder term $\bar{h}(\varepsilon s, z)$ satisfies

$$
\partial_{s}^{(i)} \bar{h}(\varepsilon s, z)=O\left(\varepsilon^{i} z^{i+5}\right) ; \quad i \geqq 0
$$

Therefore, using the same notation $\bar{h}(\varepsilon s, z)$ for a remainder term similar to the previous one we also have

$$
\frac{\partial_{z} g_{s s}}{2 g_{s s}}=-\varepsilon k-\varepsilon^{2} k^{2} z-\varepsilon^{3} k^{3} z^{2}-\varepsilon^{4} k^{4} z^{3}-\varepsilon^{5} k^{5} z^{4}+\varepsilon^{6} \bar{h}(\varepsilon s, z)
$$

Now we Taylor-expand in $\varepsilon$ the following quantities:

$$
\begin{aligned}
\frac{1}{g_{s s}}= & \frac{1}{(1-\varepsilon z k)^{2}}=1+2 \varepsilon z k+3 \varepsilon^{2} z^{2} k^{2}+\varepsilon^{3} \bar{a}(\varepsilon s, z) \\
& -\frac{\partial_{s} g_{s s}}{2 g_{s s}^{2}}=\frac{\varepsilon^{2} z k^{\prime}}{(1-\varepsilon z k)^{3}}=\varepsilon^{2} z k^{\prime}\left(1+3 \varepsilon z k+\varepsilon^{2} \bar{b}(\varepsilon s, z)\right)
\end{aligned}
$$

where the remainders $\bar{a}(\varepsilon s, z), \bar{b}(\varepsilon s, z)$ satisfy

$$
\partial_{s}^{(i)} \bar{a}(\varepsilon s, z)=O\left(\varepsilon^{i} z^{i+3}\right), \quad \partial_{s}^{(i)} \bar{b}(\varepsilon s, z)=O\left(\varepsilon^{i} z^{i+2}\right), \quad i \geqq 0
$$

In conclusion, the expansion of the Laplacian in the above coordinates $(s, z)$ (see (19)) is

$$
\begin{align*}
\Delta= & \partial_{z}^{2}+\partial_{s}^{2}-\varepsilon k \partial_{z}-\varepsilon^{2} k^{2} z \partial_{z}-\varepsilon^{3} k^{3} z^{2} \partial_{z}-\varepsilon^{4} k^{4} z^{3} \partial_{z}-\varepsilon^{5} k^{5} z^{4} \partial_{z}+\varepsilon^{6} \bar{h}(\varepsilon s, z) \\
& +\varepsilon z\left(2 k \partial_{s}^{2}+\varepsilon k^{\prime} \partial_{s}\right)+\varepsilon^{2} z^{2}\left(3 k^{2} \partial_{s}^{2}+3 \varepsilon k k^{\prime} \partial_{s}\right)+\varepsilon^{3}\left(\bar{a}(\varepsilon s, z) \partial_{s}^{2}\right. \\
& \left.+\varepsilon z k^{\prime} \bar{b}(\varepsilon s, z) \partial_{s}\right) \tag{21}
\end{align*}
$$

where, we recall, $k$ and its derivatives are evaluated at $\varepsilon s$.
Given a function

$$
f: \mathbb{R}^{2} \rightarrow \mathbb{R}
$$

of class $C^{2}$, it is possible to make the change of variables $t:=z-\phi(\varepsilon s)$. In other words, we define

$$
\tilde{f}: \mathbb{R}^{2} \rightarrow \mathbb{R}
$$

by setting $\tilde{f}(s, z):=f(s, z-\phi(\varepsilon s))$. A straightforward computation shows that

$$
\begin{aligned}
\partial_{z} \tilde{f}(s, z)= & \partial_{t} f(s, z-\phi) \\
\partial_{s} \tilde{f}(s, z)= & \partial_{s} f(y, z-\phi)-\varepsilon \phi^{\prime} \partial_{t} f(s, z-\phi) \\
\partial_{s}^{2} \tilde{f}(s, z)= & \partial_{s}^{2} f(y, z-\phi)-2 \varepsilon \phi^{\prime} \partial_{s t} f(y, z-\phi) \\
& -\varepsilon^{2} \phi^{\prime \prime} \partial_{t} f(y, z-\phi)+\varepsilon\left(\phi^{\prime}\right)^{2} \partial_{t}^{2} f(y, z-\phi)
\end{aligned}
$$

where, we recall, $\phi$ and its derivatives are evaluated at $\varepsilon s$. Hence, by (20), the expansion we are interested in, using the latter coordinates ( $s, t$ ), is given by

$$
\begin{align*}
\Delta= & \partial_{s}^{2}+\partial_{t}^{2}-\varepsilon k \partial_{t}-\varepsilon^{2}(t+\phi) k^{2} \partial_{t}-\varepsilon^{3}(t+\phi)^{2} k^{3} \partial_{t}-\varepsilon^{4}(t+\phi)^{3} k^{4} \partial_{t} \\
& -\varepsilon^{5}(t+\phi)^{4} k^{5} \partial_{t}-\varepsilon^{6} \bar{h} \partial_{t}-\varepsilon^{2} \phi^{\prime \prime} \partial_{t}-2 \varepsilon \phi^{\prime} \partial_{s t}+\varepsilon^{2}\left(\phi^{\prime}\right)^{2} \partial_{t}^{2} \\
& +\varepsilon(t+\phi)\left\{2 k \partial_{s}^{2}+\varepsilon k^{\prime} \partial_{s}-\varepsilon^{2}\left(2 k \phi^{\prime \prime}+k^{\prime} \phi^{\prime}\right) \partial_{t}-4 \varepsilon k \phi^{\prime} \partial_{s t}+2 \varepsilon^{2} k\left(\phi^{\prime}\right)^{2} \partial_{t}^{2}\right\} \\
& +\varepsilon^{2}(t+\phi)^{2}\left\{3 k^{2} \partial_{s}^{2}+3 \varepsilon k k^{\prime} \partial_{s}-\varepsilon^{2}\left(3 k^{2} \phi^{\prime \prime}+3 k k^{\prime} \phi^{\prime}\right) \partial_{t}-6 \varepsilon k^{2} \phi^{\prime} \partial_{s t}\right. \\
& \left.+3 \varepsilon^{2} k^{2}\left(\phi^{\prime}\right)^{2} \partial_{t}^{2}\right\}+\varepsilon^{3}\left\{\bar{a}(\varepsilon s, t)\left(\partial_{s}^{2}-2 \varepsilon \phi^{\prime} \partial_{s t}+\varepsilon^{2} \phi^{\prime \prime} \partial_{t}+\varepsilon\left(\phi^{\prime}\right)^{2} \partial_{t}^{2}\right)\right. \\
& \left.+\varepsilon(t+\phi) k^{\prime} \bar{b}(\varepsilon s, t)\left(\partial_{s}-\varepsilon \phi^{\prime} \partial_{t}\right)\right\}, \tag{22}
\end{align*}
$$

see also formulas (28) and (29) in [35].

### 3.3. Construction of the Approximate Solution

We proceed by fixing a function $\phi \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$ such that $\|\phi\|_{C^{4, \alpha}(\mathbb{R})}<1$ (recall (13)) and constructing an approximate solution $v_{\varepsilon, \phi}$ whose nodal set is a perturbation of the initial curve $\gamma_{T}$, tilting it transversally in the normal direction by $\phi$ (scaling properly its argument). This approximate solution is constructed in such a way that $v_{\varepsilon, \phi} \rightarrow \pm 1$ when the distance from $\gamma_{T}$ tends to infinity from different sides. More precisely, we observe that $\gamma_{T}$ divides $\mathbb{R}^{2}$ into two open unbounded regions: an upper part $\gamma_{T}^{+}$and a lower part $\gamma_{T}^{-}$.

We set

$$
\mathbb{H}(x):= \begin{cases}1 & \text { if } x \in \gamma_{T}^{+} \\ 0 & \text { if } x \in \gamma_{T} \\ -1 & \text { if } x \in \gamma_{T}^{-}\end{cases}
$$

and introduce a $C^{\infty}$ cutoff function $\zeta: \mathbb{R} \rightarrow \mathbb{R}$ such that

$$
\zeta(t)= \begin{cases}1 & \text { for } t<1 \\ 0 & \text { for } t>2\end{cases}
$$

For any $\varepsilon>0$ and for any integer $l>0$, recalling the definition of $V_{\varepsilon}$ in (18), we set

$$
\chi_{l}(x):= \begin{cases}\zeta\left(|t|-\frac{1}{8 \varepsilon}-l\right) & \text { if } x=Z_{\varepsilon}(s, t) \in V_{\varepsilon} \\ 0 & \text { if } x \in \mathbb{R}^{2} \backslash V_{\varepsilon}\end{cases}
$$

We will start by constructing an approximate solution $\hat{v}_{\varepsilon, \phi}$ in $V_{\varepsilon}$, and then globalize it using the above cut-off functions, introducing

$$
\begin{equation*}
v_{\varepsilon, \phi}(x)=\chi_{5}(x) \hat{v}_{\varepsilon, \phi}(x)+\left(1-\chi_{5}(x)\right) \mathbb{H}(x), \quad x \in \mathbb{R}^{2} . \tag{23}
\end{equation*}
$$

Since $v_{\varepsilon, \phi}$ coincides with $\hat{v}_{\varepsilon, \phi}$ near $\gamma_{T}$, it is convenient to define $\hat{v}_{\varepsilon, \phi}$ through the Fermi coordinates $(s, t)$, see (17). In order to do so, we first define a function $\tilde{v}_{\varepsilon, \phi}(s, t)$ on $\mathbb{R}^{2}$, in such a way that its zero set is close to $\{t=0\}$, then we set

$$
\hat{v}_{\varepsilon, \phi}(x):= \begin{cases}\tilde{v}_{\varepsilon, \phi}\left(Z_{\varepsilon}^{-1}(x)\right) & \text { if } x \in V_{\varepsilon} \\ 0 & \text { if } x \in \mathbb{R}^{2} \backslash V_{\varepsilon}\end{cases}
$$

In order to have a global definition of $v_{\varepsilon, \phi}$, the value of $\hat{v}_{\varepsilon, \phi}$ far from the curve is not relevant, since it is multiplied by a cut-off function that is identically zero there. We stress that, by the symmetries of $\phi, v_{\varepsilon, \phi}$ satisfies the symmetry properties (10) if $\hat{v}_{\varepsilon, \phi}$ does.

A natural first guess for an approximate solution is $\tilde{v}_{\varepsilon, \phi}(s, t):=v_{0}(t)$, where $v_{0}$ is the unique solution to the problem

$$
\begin{cases}-v_{0}^{\prime \prime}=v_{0}-v_{0}^{3} & \text { on } \mathbb{R} \\ v_{0}(0)=0 & \text { as } t \rightarrow \pm \infty \\ v_{0} \rightarrow \pm 1\end{cases}
$$

with explicit formula $v_{0}(t):=\tanh (t / \sqrt{2})$. In this way, the nodal set would be exactly the image of the curve

$$
\gamma_{T, \phi}:=\left\{\gamma_{T}(s)+\phi(\varepsilon s) \gamma_{T}^{\prime}(s)^{\perp}\right\}
$$

However, this simple approximation is not suitable for our purposes, and we need to correct it in two aspects. First, in order to recognize the linearized Willmore equation after the Lyapunov-Schmidt reduction, we have to improve the accuracy of the solution by adding further correction terms. Secondly, formally expanding in $\varepsilon$ the Cahn-Hilliard-Willmore equation on the above function we will produce error terms involving derivatives of $\phi$ up to the order six multiplied by high powers of $\varepsilon$; hence we will get an equation that, in principle, we would not be able to solve in $\phi$. In order to avoid this problem, we use a family $\left\{R_{\theta}\right\}_{\theta \geqq 1}$ of smoothing operators on periodic functions on $[0, \bar{T}]$, introduced by Alinhac and Gérard (see [4]), namely operators satisfying

$$
\begin{array}{rlr}
\left\|R_{\theta} \phi\right\|_{C^{k, \alpha}([0, \bar{T}])} \leq c\|\phi\|_{C^{k^{\prime}, \alpha^{\prime}}([0, \bar{T}])} & \text { if } k+\alpha \leqq k^{\prime}+\alpha^{\prime} ; \\
\left\|R_{\theta} \phi\right\|_{C^{k, \alpha}([0, \bar{T}])} \leq c \theta^{k+\alpha-k^{\prime}-\alpha^{\prime}}\|\phi\|_{C^{k^{\prime}, \alpha^{\prime}}([0, \bar{T}])} & \text { if } k+\alpha \geqq k^{\prime}+\alpha^{\prime} ; \\
\left\|\phi-R_{\theta} \phi\right\|_{C^{k, \alpha}([0, \bar{T}])} \leq c \theta^{k+\alpha-k^{\prime}-\alpha^{\prime}}\|\phi\|_{C^{k^{\prime}, \alpha^{\prime}}([0, \bar{T}])} & \text { if } k+\alpha \leqq k^{\prime}+\alpha^{\prime} \tag{26}
\end{array}
$$

Such operators are obtained roughly, by, truncating the Fourier modes higher than $\theta$. It is possible to find further details in [15], where the periodic case is specifically treated. This latter issue is common to interface constructions, see for example [34] and [35], and is treated in a similar manner.

Now we set $\phi_{\star}:=R_{1 / \varepsilon} \phi$ and we consider the change of variables $t:=z-$ $\phi_{\star}(\varepsilon s)$, which corresponds to replacing $\phi$ by $\phi_{\star}$ in the expansion of the Laplacian (22). The Cahn-Hilliard-Willmore equation evaluated on the above function $v_{0}(t)$ is formally of order $\varepsilon^{2}$. To correct the terms of order $\varepsilon^{2}$ we can consider

$$
v_{0}(t)+v_{1, \varepsilon, \phi}(s, t)+v_{2, \varepsilon, \phi}(s, t)
$$

as an approximate solution near the curve, as in [35, Subsection 5.1]. Here, $v_{1, \varepsilon, \phi}$ is given by

$$
\begin{equation*}
v_{1, \varepsilon, \phi}(s, t):=v_{0}\left(t+\phi_{\star}(\varepsilon s)-\phi(\varepsilon s)\right)-v_{0}(t) \tag{27}
\end{equation*}
$$

and

$$
\begin{equation*}
v_{2, \varepsilon, \phi}(s, t):=\varepsilon^{2}\left(-k^{2}(\varepsilon s)+\varepsilon L \phi_{\star}(\varepsilon s)\right) \eta(t)+\varepsilon^{2}\left(\phi_{\star}^{\prime}(\varepsilon s)\right)^{2} \tilde{\eta}(t) \tag{28}
\end{equation*}
$$

where

$$
\begin{equation*}
L \phi:=-2 k \phi^{\prime \prime}-2 k^{3} \phi \tag{29}
\end{equation*}
$$

and

$$
\eta(t)=-v_{0}^{\prime}(t) \int_{0}^{t}\left(v_{0}^{\prime}(s)\right)^{-2} \mathrm{~d} s \int_{-\infty}^{s} \frac{\tau\left(v_{0}^{\prime}(\tau)\right)^{2}}{2} \mathrm{~d} \tau
$$

In the above formulas, $v_{0}+v_{1, \varepsilon, \phi}$ represents the tilted profile $v_{0}\left(t+\phi_{\star}(\varepsilon s)-\phi(\varepsilon s)\right)$ rewritten conveniently in this way for computational purposes. The function $\eta$ is exponentially decaying, odd in $t$ and solves

$$
\begin{aligned}
L_{\star} \eta(t) & :=-\eta^{\prime \prime}(t)+W^{\prime \prime}\left(v_{0}(t)\right) \eta(t)=\frac{1}{2} t v_{0}^{\prime}(t) \\
\int_{\mathbb{R}} \eta(t) v_{0}^{\prime}(t) \mathrm{d} t & =0
\end{aligned}
$$

Here $L_{\star}$ represents the second variation of the one-dimensional Allen-Cahn energy evaluated at $v_{0}$. Similarly, $\tilde{\eta}(t):=-t v_{0}^{\prime}(t) / 2$ solves

$$
\begin{aligned}
L_{\star} \tilde{\eta}(t) & =v_{0}^{\prime \prime}(t) \\
\int_{\mathbb{R}} \tilde{\eta}(t) v_{0}^{\prime}(t) \mathrm{d} t & =0 .
\end{aligned}
$$

We note that, in particular, $L_{\star}^{2} \eta=-v_{0}^{\prime \prime}$.
In order to understand the $x_{2}$-dependence of $u_{T}$, see (5), we are interested in determining the main-order term of $\phi$, which will turn out to be of order $\varepsilon$. We then take $\phi$ of the form

$$
\phi:=\varepsilon(h+\psi)
$$

where $h$ is an explicit multiple of $\bar{\phi}$ (see Propositions 4 and 12) and $\psi$ is some fixed small $C_{\bar{T}}^{4, \alpha}(\mathbb{R})$-function, in the sense that $\|\psi\|_{C^{4, \alpha}(\mathbb{R})}<c \varepsilon$, for some constant $c>1$ to be determined later with the aid of a fixed point argument. Now we have to compute the error, that is we have to apply the Cahn-Hilliard-Willmore operator F

$$
\begin{equation*}
F(u)=-\Delta\left(-\Delta u+W^{\prime}(u)\right)+W^{\prime \prime}(u)\left(-\Delta u+W^{\prime}(u)\right) \tag{30}
\end{equation*}
$$

to the approximate solution. Since the approximate solution is defined in a neighbourhood of the perturbed curve $\gamma_{T, \phi}$, namely in $V_{\varepsilon, \phi}$, all the computations below will be performed in the coordinates $(s, t) \in \mathbb{R} \times(-1 / 4 \varepsilon, 1 / 4 \varepsilon)$. Using then cut-off functions, we then extend $F$ to be identically zero for $|t| \geq 1 / 4 \varepsilon$.

It turns out that, thanks to this last choice of the approximate solution, the Willmore equation, (22) and a Taylor expansion of the potential $W$ in (2), the error is of order $\varepsilon^{3}$. More precisely, for $|t|<1 / 4 \varepsilon$ we have

$$
\begin{align*}
F & \left(v_{0}+v_{1, \varepsilon, \phi}+v_{2, \varepsilon, \phi}\right) \\
= & -\frac{3}{2} \varepsilon^{3} k^{3}\left(2 t v_{0}^{\prime \prime}+v_{0}^{\prime}\right)+\varepsilon^{4}\left\{-4 k^{4} t^{2} v_{0}^{\prime \prime}+k^{4} \eta^{\prime \prime}-t v_{0}^{\prime}\left(3 k^{4}+\left(k^{\prime}\right)^{2}\right)\right\} \\
& +\varepsilon^{5}\left\{-\left(h_{\star}^{(4)}+\psi_{\star}^{(4)}\right) v_{0}^{\prime}+\left(h_{\star}^{\prime \prime}+\psi_{\star}^{\prime \prime}\right)\left(3 k^{2} t v_{0}^{\prime \prime}-k^{2} v_{0}^{\prime}+6 k^{2}\left(v_{0}^{\prime}+2 t v_{0}^{\prime \prime}\right)\right)\right. \\
& +\left(h_{\star}^{\prime}+\psi_{\star}^{\prime}\right) k k^{\prime}\left(8 t v_{0}^{\prime \prime}-v_{0}^{\prime}+6\left(v_{0}^{\prime}+2 t v_{0}^{\prime \prime}\right)\right) \\
& +\left(h_{\star}+\psi_{\star}\right)\left(-9 k^{4} t v_{0}^{\prime \prime}-4 k^{4} v_{0}^{\prime}-3\left(k^{\prime}\right)^{2} v_{0}^{\prime}\right) \\
& +\left(h_{\star}-h+\psi_{\star}-\psi\right)^{(4)} v_{0}^{\prime}+2 k\left(h_{\star}^{\prime}+\psi_{\star}^{\prime}\right)^{2}\left(2 v_{0}^{\prime \prime \prime}-t L_{\star} v_{0}^{\prime \prime}\right) \\
& +k^{5}\left(-4 t^{2} v_{0}^{\prime}-5 t^{3} v_{0}^{\prime \prime}+2 t \eta^{\prime \prime}+12 \eta \eta^{\prime} v_{0}+6 \eta^{2} v_{0}^{\prime}+\frac{5}{2} \eta^{\prime}\right) \\
& \left.+k\left(k^{\prime}\right)^{2}\left(-9 t^{2} v_{0}^{\prime}-4 \eta^{\prime}\right)\right\}+\varepsilon^{6} F_{\varepsilon}^{1}(\psi) . \tag{31}
\end{align*}
$$

In (31), the right-hand side is evaluated at $(\varepsilon s, t)$. The term $F_{\varepsilon}^{1}$ is defined to be identically zero for $|t| \geq 1 / 4 \varepsilon$ and can be suitably estimated using weighted norms. To introduce these, for any $0<\delta<\sqrt{2}$ and $x=\left(x_{1}, x_{2}\right) \in \mathbb{R}^{2}$ define

$$
\begin{equation*}
\varphi_{\varepsilon, \delta}(x):=\int_{\gamma_{T}} G_{\delta}(x, y) \mathrm{d} l(y) \tag{32}
\end{equation*}
$$

where $G_{\delta}$ is the Green function of $-\Delta+\delta^{2}$ in $\mathbb{R}^{2}$. We denote by $C^{n, \alpha}\left(\mathbb{R}^{2}\right)$ the space of functions $u: \mathbb{R}^{2} \rightarrow \mathbb{R}$ that are $n$ times differentiable and whose $n$th derivatives are Hölder continuous with exponent $\alpha$. For $L:=\left(\gamma_{T}\right)_{1}(T)-\left(\gamma_{T}\right)_{1}(0)$, we set

$$
\begin{aligned}
C_{L, \delta}^{n, \alpha}\left(\mathbb{R}^{2}\right) & :=\left\{u \in C^{n, \alpha}\left(\mathbb{R}^{2}\right):\left\|u \varphi_{\varepsilon, \delta}^{-1}\right\|_{C^{n, \alpha}\left(\mathbb{R}^{2}\right)}<\infty, u\left(x_{1}, x_{2}\right)=-u\left(-x_{1},-x_{2}\right)\right. \\
& \left.=-u\left(x_{1}+\frac{L}{2},-x_{2}\right)\right\}
\end{aligned}
$$

where

$$
\begin{equation*}
\|u\|_{C^{n, \alpha}\left(\mathbb{R}^{2}\right)}=\sum_{j=0}^{n}\left\|\nabla^{j} u\right\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}+\sup _{x \neq y|\beta|=n} \sup _{\left|\partial_{\beta} u(x)-\partial_{\beta} u(y)\right|}^{|x-y|^{\alpha}} \tag{33}
\end{equation*}
$$

Functions belonging to these spaces decay exponentially away from the curve $\gamma_{T}$, with rate $\mathrm{e}^{-\delta d\left(\cdot, \gamma_{T}\right)}$, and satisfy the symmetries of the curve, that is they are even, periodic with period $L$, and they change sign after translation of half a period and a reflection about the $x_{2}$ axis. We endow these spaces with the norms

$$
\begin{equation*}
\|u\|_{C_{\delta}^{n, \alpha}\left(\mathbb{R}^{2}\right)}:=\left\|u \varphi_{\varepsilon, \delta}^{-1}\right\|_{C^{n, \alpha}\left(\mathbb{R}^{2}\right)} \tag{34}
\end{equation*}
$$

Using this notation and recalling (23), we have that the error term $F_{\varepsilon}^{1}$ in (31) satisfies

$$
\left\{\begin{array}{l}
\left\|\chi_{4} F_{\varepsilon}^{1}(\psi)\right\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c  \tag{35}\\
\left\|\chi_{4} F_{\varepsilon}^{1}\left(\psi_{1}\right)-\chi_{4} F_{\varepsilon}^{2}\left(\psi_{2}\right)\right\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c\left\|\psi_{1}-\psi_{2}\right\|_{C^{4, \alpha}(\mathbb{R})}
\end{array}\right.
$$

for $\psi, \psi_{1}, \psi_{2} \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$ such that $\|\psi\|_{C^{4, \alpha}(\mathbb{R})},\left\|\psi_{i}\right\|_{C^{4, \alpha}(\mathbb{R})}<1, i=1,2$.
In what follows, we will solve the Cahn-Hilliard-Willmore equation through a Lyapunov-Schmidt reduction, and to deal with the bifurcation equation we will need to consider the projection of the error terms in the Cahn-Hilliard-Willmore equation along the kernel of its linearised operator. Fixing $s$, this corresponds to multiplying the error term by $v_{0}^{\prime}(t)$ (and a cut-off function in $t$ ) and integrating in $t$. For instance, if $\chi_{l}$ is the cut-off function introduced at the beginning of this subsection, the projection

$$
G_{\varepsilon}^{1}(\psi)(s):=\int_{\mathbb{R}} \chi_{4}(t) F_{\varepsilon}^{1}(\psi)(s, t) v_{0}^{\prime}(t) \mathrm{d} t
$$

of $\chi_{4} F_{\varepsilon}^{1}(\psi)$ fulfils

$$
\left\{\begin{array}{l}
\left\|G_{\varepsilon}^{1}(\psi)\right\|_{C^{0, \alpha}(\mathbb{R})} \leqq c  \tag{36}\\
\left\|G_{\varepsilon}^{1}\left(\psi_{1}\right)-G_{\varepsilon}^{1}\left(\psi_{2}\right)\right\|_{C^{0, \alpha}(\mathbb{R})} \leqq c\left\|\psi_{1}-\psi_{2}\right\|_{C^{4, \alpha}(\mathbb{R})}
\end{array}\right.
$$

In other words, apart from the coefficient of order $\varepsilon^{6}$, we get a remainder which is uniformly bounded for $\psi$ in the unit ball of $C_{\bar{T}}^{4, \alpha}(\mathbb{R})$, with Lipschitz dependence.

Setting

$$
\begin{equation*}
c_{\star}:=\int_{\mathbb{R}}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t>0 \tag{37}
\end{equation*}
$$

and using the fact that

$$
\int_{\mathbb{R}} t v_{0}^{\prime \prime} v_{0}^{\prime} \mathrm{d} t=-\frac{1}{2} c_{\star},
$$

we can see that the projection of the linear term in $h+\psi$ (and their derivatives) appearing at order $\varepsilon^{5}$ is given by

$$
\begin{align*}
\int_{\mathbb{R}}\{ & -\left(h_{\star}^{(4)}+\psi_{\star}^{(4)}\right) v_{0}^{\prime}+\left(h_{\star}^{\prime \prime}+\psi_{\star}^{\prime \prime}\right)\left(3 k^{2} t v_{0}^{\prime \prime}-k^{2} v_{0}^{\prime}+6 k^{2}\left(v_{0}^{\prime}+2 t v_{0}^{\prime \prime}\right)\right) \\
& +\left(h_{\star}^{\prime}+\psi_{\star}^{\prime}\right) k k^{\prime}\left(8 t v_{0}^{\prime \prime}-v_{0}^{\prime}+6\left(v_{0}^{\prime}+2 t v_{0}^{\prime \prime}\right)\right) \\
& \left.+\left(h_{\star}+\psi_{\star}\right)\left(-9 k^{4} t v_{0}^{\prime \prime}-4 k^{4} v_{0}^{\prime}-3\left(k^{\prime}\right)^{2} v_{0}^{\prime}\right)+\left(h_{\star}-h+\psi_{\star}-\psi\right)^{(4)} v_{0}^{\prime}\right\} v_{0}^{\prime} \mathrm{d} t \\
= & -c_{\star}\left(\tilde{L}_{0}\left(h_{\star}+\psi_{\star}\right)-\left(h-h_{\star}+\psi-\psi_{\star}\right)^{(4)}\right)= \\
& -c_{\star}\left(\tilde{L}_{0}(h+\psi)+\left(\tilde{L}_{0}-\frac{\mathrm{d}^{4}}{\mathrm{~d} s^{4}}\right)\left(h_{\star}-h+\psi_{\star}-\psi\right)\right) \tag{38}
\end{align*}
$$

where we recall that (see (12))

$$
\tilde{L}_{0} \phi:=\phi^{(4)}+\frac{5}{2}\left(k^{2} \phi^{\prime}\right)^{\prime}+\left(3\left(k^{\prime}\right)^{2}-\frac{1}{2} k^{4}\right) \phi
$$

The terms of order $\varepsilon^{3}$ in (31) can be eliminated by adding to the approximate solution an extra correction of the form

$$
v_{3, \varepsilon, \phi}(s, t):=\frac{3}{2} \varepsilon^{3} k^{3} \eta_{1}
$$

where $\eta_{1}$ solves

$$
\left\{\begin{array}{l}
L_{\star}^{2} \eta_{1}=2 t v_{0}^{\prime \prime}+v_{0}^{\prime}  \tag{39}\\
\int_{\mathbb{R}} \eta_{1} v_{0}^{\prime} \mathrm{d} t=0
\end{array}\right.
$$

We point out that (39) is solvable since the first right-hand side is orthogonal to $v_{0}^{\prime}$, that is

$$
\int_{\mathbb{R}}\left(2 t v_{0}^{\prime \prime}+v_{0}^{\prime}\right) v_{0}^{\prime} \mathrm{d} t=0
$$

As it is well-known, see example [28], the (decaying) kernel of $L_{\star}$ is generated by $v_{0}^{\prime}$ so the existence of $\eta_{1}$ follows from Fredholm's theory.

As a consequence, using (22) once again and an expansion similar to (31), for $|t|<1 / 4 \varepsilon$ we have

$$
\begin{align*}
F & \left(v_{0}+v_{1, \varepsilon, \phi}+v_{2, \varepsilon, \phi}+v_{3, \varepsilon, \phi}\right)=\varepsilon^{4}\left\{-4 k^{4} t^{2} v_{0}^{\prime \prime}+k^{4} \eta^{\prime \prime}\right. \\
& \left.-t v_{0}^{\prime}\left(3 k^{4}+\left(k^{\prime}\right)^{2}\right)+3 k^{4}\left(L_{\star} \eta_{1}\right)^{\prime}\right\} \\
& +\varepsilon^{5}\left\{\left(h-h_{\star}+\psi_{\star}-\psi\right)^{(4)} v_{0}^{\prime}-\left(h^{(4)}+\psi_{\star}^{(4)}\right) v_{0}^{\prime}\right. \\
& +\left(h^{\prime \prime}+\phi_{\star}^{\prime \prime}\right)\left(3 k^{2} t v_{0}^{\prime \prime}-k^{2} v_{0}^{\prime}+6 k^{2}\left(v_{0}^{\prime}+2 t v_{0}^{\prime \prime}\right)\right) \\
& +\left(h_{\star}^{\prime}+\psi_{\star}^{\prime}\right) k k^{\prime}\left(8 t v_{0}^{\prime \prime}-v_{0}^{\prime}+6\left(v_{0}^{\prime}+2 t v_{0}^{\prime \prime}\right)\right) \\
& +\left(h+\psi_{\star}\right)\left(-9 k^{4} t v_{0}^{\prime \prime}-4 k^{4} v_{0}^{\prime}-3\left(k^{\prime}\right)^{2} v_{0}^{\prime}\right) \\
& +k^{5}\left(-4 t^{2} v_{0}^{\prime}-5 t^{3} v_{0}^{\prime \prime}+2 t \eta^{\prime \prime}+12 \eta \eta^{\prime} v_{0}+6 \eta^{2} v_{0}^{\prime}+\frac{5}{2} \eta^{\prime}\right. \\
& \left.+3 t\left(L_{\star} \eta_{1}\right)^{\prime}-\frac{3}{2} \eta_{1}^{\prime \prime}+\frac{9}{2} L_{\star} \eta_{1}\right) \\
& \left.+k\left(k^{\prime}\right)^{2}\left(-9 t^{2} v_{0}^{\prime}-4 \eta^{\prime}-18 L_{\star} \eta_{1}\right)+2 k\left(h_{\star}^{\prime}+\psi_{\star}^{\prime}\right)^{2}\left(2 v_{0}^{\prime \prime \prime}-t L_{\star} v_{0}^{\prime \prime}\right)\right\} \\
& +\varepsilon^{6}\left\{F_{\varepsilon}^{1}(\psi)+F_{\varepsilon}^{2}(\psi)\right\} . \tag{40}
\end{align*}
$$

Notice that $F\left(v_{0}+v_{1, \varepsilon, \phi}+v_{2, \varepsilon, \phi}+v_{3, \varepsilon, \phi}\right)$ and $F_{\varepsilon}^{2}(\psi)$ vanish identically for $|t| \geq 1 / 4 \varepsilon, F_{\varepsilon}^{2}(\psi)$ satisfies (35) and the projection

$$
G_{\varepsilon}^{2}(\psi)(s):=\int_{\mathbb{R}} \chi_{4}(t) F_{\varepsilon}^{2}(\psi)(s, t) v_{0}^{\prime}(t) \mathrm{d} t
$$

fulfils estimates similar to (36). Once again, in (40) the right-hand side is evaluated at ( $\varepsilon s, t$ ). Similarly, we can improve our approximate solution by correcting the terms of order $\varepsilon^{4}$ in (40). Integrating by parts we have that

$$
\int_{\mathbb{R}}\left\{-3\left(L_{\star} \eta_{1}\right)^{\prime}+4 t^{2} v_{0}^{\prime \prime}+3 t v_{0}^{\prime}-\eta^{\prime \prime}-\frac{1}{2} Q(\eta, \eta)\right\} v_{0}^{\prime} \mathrm{d} t=\int_{\mathbb{R}} t\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t=0
$$

where we have set

$$
\begin{align*}
Q(v, w) & :=-\left(W^{\prime \prime \prime}\left(v_{0}\right) v w\right)^{\prime \prime}+W^{\prime \prime}\left(v_{0}\right) W^{\prime \prime \prime}\left(v_{0}\right) v w+W^{\prime \prime \prime}\left(v_{0}\right)\left(v L_{\star} w+w L_{\star} v\right) \\
& =L_{\star}\left(W^{\prime \prime \prime}\left(v_{0}\right) v w\right)+W^{\prime \prime \prime}\left(v_{0}\right)\left(v L_{\star} w+w L_{\star} v\right) . \tag{41}
\end{align*}
$$

Therefore by the above comments we can solve

$$
\left\{\begin{array}{l}
L_{\star}^{2} \eta_{2}=-3\left(L_{\star} \eta_{1}\right)^{\prime}+4 t^{2} v_{0}^{\prime \prime}+3 t v_{0}^{\prime}-\eta^{\prime \prime}-\frac{1}{2} Q(\eta, \eta) \\
\int_{\mathbb{R}} \eta_{2} v_{0}^{\prime} \mathrm{d} t=0
\end{array}\right.
$$

and

$$
\left\{\begin{array}{l}
L_{\star}^{2} \eta_{3}=t v_{0}^{\prime} \\
\int_{\mathbb{R}} \eta_{3} v_{0}^{\prime} \mathrm{d} t=0
\end{array}\right.
$$

Finally, we set

$$
\begin{equation*}
\tilde{v}_{\varepsilon, \phi}(s, t):=v_{0}(t)+v_{1, \varepsilon, \phi}(s, t)+v_{2, \varepsilon, \phi}(s, t)+v_{3, \varepsilon, \phi}(s, t)+v_{4, \varepsilon, \phi}(s, t), \tag{42}
\end{equation*}
$$

where $v_{4, \varepsilon, \phi}(s, t)=\varepsilon^{4}\left(k^{4} \eta_{2}+\left(k^{\prime}\right)^{2} \eta_{3}\right)$. For $|t|<1 / 4 \varepsilon$, using expansions similar to the previous ones we compute

$$
\begin{align*}
& F\left(\tilde{v}_{\varepsilon, \phi}\right)=\varepsilon^{5}\left\{-\left(h_{\star}^{(4)}+\psi_{\star}^{(4)}\right) v_{0}^{\prime}+\left(h_{\star}^{\prime \prime}+\psi_{\star}^{\prime \prime}\right)\left(3 k^{2} t v_{0}^{\prime \prime}-k^{2} v_{0}^{\prime}+6 k^{2}\left(v_{0}^{\prime}+2 t v_{0}^{\prime \prime}\right)\right)\right. \\
& \quad+\left(h_{\star}^{\prime}+\psi_{\star}^{\prime}\right) k k^{\prime}\left(8 t v_{0}^{\prime \prime}-v_{0}^{\prime}+6\left(v_{0}^{\prime}+2 t v_{0}^{\prime \prime}\right)\right) \\
& \quad+\left(h_{\star}+\psi_{\star}\right)\left(-9 k^{4} t v_{0}^{\prime \prime}-4 k^{4} v_{0}^{\prime}-3\left(k^{\prime}\right)^{2} v_{0}^{\prime}\right)+\left(h_{\star}-h+\phi_{\star}-\phi\right)^{(4)} v_{0}^{\prime} \\
& \left.\quad+E_{5}(\varepsilon s, t)+2 k\left(h_{\star}^{\prime}+\psi_{\star}^{\prime}\right)^{2}\left(2 v_{0}^{\prime \prime \prime}-t L_{\star} v_{0}^{\prime \prime}\right)\right\}+\varepsilon^{6}\left\{F_{\varepsilon}^{1}(\psi)+F_{\varepsilon}^{2}(\psi)+F_{\varepsilon}^{3}(\psi)\right\} . \tag{43}
\end{align*}
$$

Here

$$
\begin{aligned}
E_{5}(s, t):= & k^{5}\left(-4 t^{2} v_{0}^{\prime}-5 t^{3} v_{0}^{\prime \prime}+2 t \eta^{\prime \prime}-\frac{3}{2} Q\left(\eta, \eta_{1}\right)+3 t\left(L_{\star} \eta_{1}\right)^{\prime}\right. \\
& \left.+12 \eta \eta^{\prime} v_{0}+6 \eta^{2} v_{0}^{\prime}-\frac{3}{2} \eta_{1}^{\prime \prime}+\frac{9}{2} L_{\star} \eta_{1}+\frac{5}{2} \eta^{\prime}+2\left(L_{\star} \eta_{2}\right)^{\prime}\right) \\
& +k\left(k^{\prime}\right)^{2}\left(2\left(L_{\star} \eta_{3}\right)^{\prime}-9 t^{2} v_{0}^{\prime}-4 \eta^{\prime}-18 L_{\star} \eta_{1}\right)
\end{aligned}
$$

$F_{\varepsilon}^{3}$ is identically zero for $|t| \geq 1 / 4 \varepsilon$ and satisfies the counterpart of (35) and

$$
G_{\varepsilon}^{3}(\psi)(s):=\int_{\mathbb{R}} \chi_{4}(t) F_{\varepsilon}^{3}(\psi)(s, t) v_{0}^{\prime}(t) \mathrm{d} t
$$

fulfils estimates similar to (36). In order to handle the error, we introduce a suitable function space and we endow it with an appropriate weighted norm. We set, for $0<\delta<\sqrt{2}$,

$$
\begin{equation*}
\psi_{\delta}(x):=\zeta\left(\left|x_{2}\right|\right)+\left(1-\zeta\left(\left|x_{2}\right|\right)\right) \mathrm{e}^{-\delta\left|x_{2}\right|} \tag{44}
\end{equation*}
$$

where $\zeta$ is defined in (3.3). Now, we define the spaces

$$
\begin{aligned}
D_{T, \delta}^{n, \alpha}\left(\mathbb{R}^{2}\right) & :=\left\{U \in C^{n, \alpha}:\left\|U \psi_{\delta}\right\|_{C^{n, \alpha}\left(\mathbb{R}^{2}\right)}<\infty, U\left(x_{1}, x_{2}\right)=-U\left(-x_{1},-x_{2}\right)\right. \\
& \left.=-U\left(x_{1}+\frac{T}{2},-x_{2}\right)\right\}
\end{aligned}
$$

endowed with the norms

$$
\begin{equation*}
\|u\|_{D_{\delta}^{n, \alpha}\left(\mathbb{R}^{2}\right)}:=\left\|u \psi_{\delta}\right\|_{C^{n, \alpha}\left(\mathbb{R}^{2}\right)} \tag{45}
\end{equation*}
$$

The difference between the space $D_{T, \delta}^{n, \alpha}\left(\mathbb{R}^{2}\right)$ and the space $C_{L, \delta}^{n, \alpha}\left(\mathbb{R}^{2}\right)$ introduced previously are the weight function, which depends just on one variable in the case of $D_{T, \delta}^{n, \alpha}\left(\mathbb{R}^{2}\right)$, and the period.

Recalling (37), define the constant

$$
\begin{equation*}
d_{\star}:=\int_{\mathbb{R}} t^{2}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t>0 \tag{46}
\end{equation*}
$$

and recall the definition of $\bar{\phi}$ in Proposition 4. From the previous computations, we have the following result:

Proposition 6. There exist a constant $\tilde{c}>0$ such that

$$
\begin{equation*}
\left\|F\left(\tilde{v}_{\varepsilon, \phi}\right)\right\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq \tilde{c} \varepsilon^{5} \tag{47}
\end{equation*}
$$

for any $\phi \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$ such that $\phi=\varepsilon\left(\frac{d_{\star}}{c_{\star}} \bar{\phi}+\psi\right)$, with $\|\psi\|_{C^{4, \alpha}(\mathbb{R})}<1$.
Remark 7. The estimate in Proposition 6 holds for any function $\phi$ of order $\varepsilon$ in $C^{4, \alpha}$ norm. However, for later purposes, we will need to take $\phi=\varepsilon\left(\frac{d_{\star}}{c_{\star}} \bar{\phi}+\psi\right)$ as above in order to determine the principal term in the expansion after projecting onto $v_{0}^{\prime}$, when dealing with the bifurcation equation.

## 4. The Lyapunov-Schmidt Reduction

Up to now, we have only constructed an approximate solution to (3), not a true solution, since $F\left(v_{\varepsilon, \phi}\right)$ is small but not zero (see (30) and (23)). Therefore we try to add a small correction $w=w_{\varepsilon, \phi}: \mathbb{R}^{2} \rightarrow \mathbb{R}$ in such a way that $F\left(v_{\varepsilon, \phi}+w\right)=0$. Rephrasing our problem in this way, the unknowns are $\phi$ and $w$, for any $\varepsilon>0$ small but fixed (recall that $\varepsilon=\frac{\bar{T}}{T}$ ). Expanding $F$ in Taylor series, our equation becomes

$$
F\left(v_{\varepsilon, \phi}\right)+F^{\prime}\left(v_{\varepsilon, \phi}\right)[w]+Q_{\varepsilon, \phi}(w)=0
$$

where

$$
\begin{equation*}
Q_{\varepsilon, \phi}(w)=\int_{0}^{1} \mathrm{~d} t \int_{0}^{t} F^{\prime \prime}\left(v_{\varepsilon, \phi}+s w\right)[w, w] \mathrm{d} s \tag{48}
\end{equation*}
$$

In order to study (4), we use a Lyapunov-Schmidt reduction, consisting in an auxiliary equation in $w$ and a bifurcation equation in $\phi$.

### 4.1. The Auxiliary Equation: A Gluing Procedure

Recalling the definition of the cut-off $\chi_{l}$ in Section 3.3, we look for a correction $w$ of the form

$$
w(x)=\chi_{2}(x) \hat{U}(x)+V(x)
$$

where $V, \hat{U}$ are defined in $\mathbb{R}^{2}$. Since $\hat{U}$ is multiplied by a cut-off function that is identically zero far from $\gamma_{T}$, we look for some suitable function $U=U(t, s)$ defined in $\mathbb{R}^{2}$, then we set (see (18))

$$
\hat{U}(x):= \begin{cases}U\left(Z_{\varepsilon}^{-1}(x)\right), & \text { if } x \in V_{\varepsilon} \\ 0 & \text { if } x \in \mathbb{R}^{2} \backslash V_{\varepsilon}\end{cases}
$$

As above, the value of $\hat{U}$ far from the curve does not matter, since it is multiplied by a cut-off function.

Remark 8. Let $\tilde{v}_{\varepsilon, \phi}$ be as in Proposition 6 and let $v_{\varepsilon, \phi}$ be as in (23) (see also the subsequent formula). Then the potential

$$
\Gamma_{\varepsilon, \phi}(x):=\left(1-\chi_{1}(x)\right) W^{\prime \prime}\left(v_{\varepsilon, \phi}\right)+\chi_{1}(x) W^{\prime \prime}(1) \quad\left(W^{\prime \prime}(1)=2\right)
$$

is positive and bounded away from 0 in the whole $\mathbb{R}^{2}$. Precisely, for any $0<\delta<$ $\sqrt{2}$, we have $0<\delta^{2}<\Gamma_{\varepsilon, \phi}(x)<2$ provided $\varepsilon$ is small enough, the estimate being uniform in $\phi$. By construction, $\Gamma_{\varepsilon, \phi} \in C^{4, \alpha}\left(\mathbb{R}^{2}\right)$, it is periodic of period $L$ (the $x_{1}$-period of $\gamma_{T}$ ), and the $L^{\infty}$ norms of the derivatives are bounded uniformly in $\phi$ and in $\varepsilon$.

Using the fact that $\chi_{2} \chi_{1}=\chi_{1}, \chi_{2} \chi_{4}=\chi_{2}($ recall (23)) and the Taylor expansion (4), we can see that the Cahn-Hilliard-Willmore equation $F\left(v_{\varepsilon, \phi}+w\right)=0$ can be rewritten as

$$
\begin{aligned}
F( & \left.v_{\varepsilon, \phi}\right)+F^{\prime}\left(v_{\varepsilon, \phi}\right) w+Q_{\varepsilon, \phi}(w) \\
= & \chi_{2}\left\{\chi_{4} F\left(\hat{v}_{\varepsilon, \phi}\right)+F^{\prime}\left(\hat{v}_{\varepsilon, \phi}\right) \hat{U}+\chi_{1} Q_{\varepsilon, \phi}\left(\chi_{2} \hat{U}+V\right)+\chi_{1} \mathrm{M}_{\varepsilon, \phi}(V)\right\} \\
& +\left(-\Delta+\Gamma_{\varepsilon, \phi}\right)^{2} V+\left(1-\chi_{2}\right) F\left(v_{\varepsilon, \phi}\right)+\left(1-\chi_{1}\right) Q_{\varepsilon, \phi}\left(\chi_{2} \hat{U}+V\right) \\
& +\mathrm{N}_{\varepsilon, \phi}(\hat{U})+\mathrm{P}_{\varepsilon, \phi}(V)
\end{aligned}
$$

where

$$
\begin{align*}
\mathrm{M}_{\varepsilon, \phi}(V):= & \left(W^{\prime \prime}\left(\hat{v}_{\varepsilon, \phi}\right)-W^{\prime \prime}(1)\right)\left(-\Delta V+\Gamma_{\varepsilon, \phi} V\right) \\
& +\left(-\Delta+W^{\prime \prime}\left(\hat{v}_{\varepsilon, \phi}\right)\right)\left[\left(W^{\prime \prime}\left(\hat{v}_{\varepsilon, \phi}\right)-W^{\prime \prime}(1)\right) V\right] ;  \tag{49}\\
\mathrm{N}_{\varepsilon, \phi}(\hat{U}):= & -2\left\langle\nabla \chi_{2}, \nabla\left(-\Delta \hat{U}+W^{\prime \prime}\left(\hat{v}_{\varepsilon, \phi}\right) \hat{U}\right)\right\rangle-\Delta \chi_{2}\left(-\Delta \hat{U}+W^{\prime \prime}\left(\hat{v}_{\varepsilon, \phi}\right) \hat{U}\right) \\
& +\left(-\Delta+W^{\prime \prime}\left(\hat{v}_{\varepsilon, \phi}\right)\right)\left(-2\left\langle\nabla \chi_{2}, \nabla \hat{U}\right\rangle-\Delta \chi_{2} \hat{U}\right) ;  \tag{50}\\
\mathrm{P}_{\varepsilon, \phi}(V):= & -2\left\langle\nabla \chi_{1}, \nabla\left(\left(W^{\prime \prime}\left(\hat{v}_{\varepsilon, \phi}\right)-W^{\prime \prime}(1)\right) V\right)\right\rangle-\Delta \chi_{1}\left(W^{\prime \prime}\left(\hat{v}_{\varepsilon, \phi}\right)-W^{\prime \prime}(1)\right) V \\
& +W^{\prime \prime \prime}\left(v_{\varepsilon, \phi}\right)\left(-\Delta v_{\varepsilon, \phi}+W^{\prime}\left(v_{\varepsilon, \phi}\right)\right) V . \tag{51}
\end{align*}
$$

By the expansion of the Laplacian (22), we can see that, expressing $F^{\prime}\left(\tilde{v}_{\varepsilon, \phi}\right)$ in the ( $s, t$ )-coordinates, for $|t|<1 / 4 \varepsilon$,

$$
\begin{equation*}
F^{\prime}\left(\tilde{v}_{\varepsilon, \phi}\right)=\mathcal{L}^{2}+\mathrm{R}_{\varepsilon, \phi} \tag{52}
\end{equation*}
$$

where

$$
\begin{equation*}
\mathcal{L}=-\left(\partial_{s}^{2}+\partial_{t}^{2}\right)+W^{\prime \prime}\left(v_{0}(t)\right) \tag{53}
\end{equation*}
$$

and $\mathrm{R}_{\varepsilon, \phi}=O(\varepsilon)$, in the sense that (recall (45))

$$
\left\|\chi_{4} \mathrm{R}_{\varepsilon, \phi} U\right\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c \varepsilon\|U\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)}
$$

Once again, we have extended $\mathrm{R}_{\varepsilon, \phi}$ to be identically zero for $|t| \geq 1 / 4 \varepsilon$. Hence we have reduced our problem to finding a solution $(\phi, V, U)$ to the system

$$
\begin{cases}\left(-\Delta+\Gamma_{\varepsilon, \phi}\right)^{2} V+\left(1-\chi_{2}\right) F\left(v_{\varepsilon, \phi}\right) &  \tag{54}\\ \quad+\left(1-\chi_{1}\right) Q_{\varepsilon, \phi}\left(\chi_{2} \hat{U}+V\right) & \text { in } \mathbb{R}^{2} \\ \quad+\mathrm{N}_{\varepsilon, \phi}(\hat{U})+\mathrm{P}_{\varepsilon, \phi}(V)=0 ; & \text { for }|t| \leqq 1 / 8 \varepsilon+4 \\ \chi_{4} F\left(\tilde{v}_{\varepsilon, \phi}\right)+\mathcal{L}^{2} U+\chi_{4} \mathrm{R}_{\varepsilon, \phi} U+\chi_{1} Q_{\varepsilon, \phi}\left(\chi_{2} U\right. & \\ \left.\quad+V\left(Z_{\varepsilon}(s, t)\right)\right)+\chi_{1} \mathrm{M}_{\varepsilon, \phi}(V)=0 & \end{cases}
$$

It is understood that, in equation (55), the cut-off functions and $V$ are evaluated at $Z_{\varepsilon}(s, t)$, see (17). First we fix $\phi$ and $U$ and we solve the auxiliary equation (54) by a fixed point argument, using the coercivity of the operator $\left(-\Delta+\Gamma_{\varepsilon, \phi}\right)^{2}$. This is possible due to fact that the potential $\Gamma_{\varepsilon, \phi}$ is bounded from above and from below by positive constants (see Remark 8).

We have next the following result, that will be proved in Section 5 (recall (45) and (34)):
Proposition 9. For any $\varepsilon>0$ small enough, for any $U \in D_{T, \delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)$ such that $\|U\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)}<1$ and for any $\phi \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$ with $\|\phi\|_{C^{4, \alpha}(\mathbb{R})}<1$, equation (54) admits a solution $V_{\varepsilon, \phi, U} \in C_{L, \delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)$ satisfying

$$
\left\{\begin{array}{l}
\left\|V_{\varepsilon, \phi, U}\right\|_{C_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c_{1} \mathrm{e}^{-\delta / 8 \varepsilon} ; \\
\left\|V_{\varepsilon, \phi, U_{1}}-V_{\varepsilon, \phi, U_{2}}\right\|_{C_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c_{1} \mathrm{e}^{-\delta / 8 \varepsilon}\left\|U_{1}-U_{2}\right\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \\
\left\|V_{\varepsilon, \phi 1, U}-V_{\varepsilon, \phi_{2}, U}\right\|_{C_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c_{1} \mathrm{e}^{-\delta / 8 \varepsilon}\left\|\phi_{1}-\phi_{2}\right\|_{C^{4, \alpha}(\mathbb{R})}
\end{array}\right.
$$

for any $U_{1}, U_{2}$ with $\left\|U_{1}\right\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)},\left\|U_{2}\right\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)}<1$, for any $\phi_{1}, \phi_{2} \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$ with $\left\|\phi_{i}\right\|_{C^{4, \alpha}(\mathbb{R})}<1, i=1,2$, and for some constant $c_{1}>0$ independent of $U, \varepsilon$ and $\phi$.

Since we reduced solving the Cahn-Hilliard-Willmore equation to the system (54)-(55), it remains to solve the second component. The operator $\mathcal{L}^{2}$ (see (53)) is not uniformly coercive as $\varepsilon \rightarrow 0$ : in fact, in the $t$ component it annihilates $v_{0}^{\prime}(t)$, while due to the fact that $s$ lies in an expanding domain, the spectrum of $\partial_{s}^{2}$ approaches zero. Due to the consequent lack of invertibility of $\mathcal{L}^{2}$ we need some orthogonality condition to solve equation $\mathcal{L}^{2} U=f$, that is

$$
\int_{\mathbb{R}} f(s, t) v_{0}^{\prime}(t) \mathrm{d} t=0 \quad \forall s \in \mathbb{R}
$$

as we will see in Section 5.1, and the solution will satisfy the same orthogonality condition (for a detailed discussion, see Section 5). As a consequence, equation (55) cannot be solved directly, through a fixed point argument, hence we subtract the projection along $v_{0}^{\prime}$ of the right-hand side. In other words, setting

$$
\begin{equation*}
\mathrm{T}(U, V, \phi):=\chi_{1} Q_{\varepsilon, \phi}\left(\chi_{2} U+V\right)+\chi_{4} \mathrm{R}_{\varepsilon, \phi}(U)+\chi_{1} \mathrm{M}_{\varepsilon, \phi}(V) \tag{56}
\end{equation*}
$$

and

$$
\begin{equation*}
p_{\phi}(s):=\frac{1}{c_{\star}} \int_{-\infty}^{\infty}\left(\chi_{4} F\left(\tilde{v}_{\varepsilon, \phi}\right)+\mathrm{T}\left(U, V_{\varepsilon, \phi, U}, \phi\right)\right)(s, t) v_{0}^{\prime}(t) \mathrm{d} t \tag{57}
\end{equation*}
$$

we can solve

$$
\begin{align*}
& \mathcal{L}^{2} U=-\chi_{4} F\left(\tilde{v}_{\varepsilon, \phi}\right)-\mathrm{T}\left(U, V_{\varepsilon, \phi, U}, \phi\right)+p_{\phi}(s) v_{0}^{\prime}(t) \\
& \int_{\mathbb{R}} U(s, t) v_{0}^{\prime}(t) \mathrm{d} t=\forall s \in \mathbb{R} \tag{58}
\end{align*}
$$

in $U$, for any small but fixed $\phi \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$. Concerning the operator near $\gamma_{T}$, we have the following result, that will be proved in Section 5:

Proposition 10. For any $\varepsilon>0$ small enough and for any $\phi \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$ with $\|\phi\|_{C^{4, \alpha}(\mathbb{R})}<1$, we can find a solution $U_{\varepsilon, \phi} \in D_{T, \delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)$ to equation (58) satisfying the orthogonality condition

$$
\begin{equation*}
\int_{\mathbb{R}} U_{\varepsilon, \phi}(s, t) v_{0}^{\prime}(t) \mathrm{d} t=0, \quad \forall s \in \mathbb{R} \tag{59}
\end{equation*}
$$

and the estimates

$$
\left\{\begin{array}{l}
\left\|U_{\varepsilon, \phi}\right\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c_{2} \varepsilon^{5}  \tag{60}\\
\left\|U_{\varepsilon, \phi_{1}}-U_{\varepsilon, \phi_{2}}\right\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c_{2} \varepsilon^{5}\left\|\phi_{1}-\phi_{2}\right\|_{C^{4, \alpha}(\mathbb{R})}
\end{array}\right.
$$

for any $\phi_{1}, \phi_{2} \in C_{T}^{4, \alpha}(\mathbb{R})$ with $\left\|\phi_{i}\right\|_{C^{4, \alpha}(\mathbb{R})}<1, i=1,2$, for some constant $c_{2}>0$ independent of $\varepsilon$.

### 4.2. The Bifurcation Equation

Using the notation in the previous subsection (see in particular the discussion before Proposition 10), the Cahn-Hilliard-Willmore equation reduces to

$$
\mathcal{L}^{2} U=-\chi_{4} F\left(\tilde{v}_{\varepsilon, \phi}\right)-\mathrm{T}\left(U, V_{\varepsilon, \phi, U}, \phi\right)
$$

Recalling (58), in order to conclude the proof it remains to solve the bifurcation equation

$$
\begin{equation*}
p_{\phi}(s)=0 \quad \text { for all } s \in \mathbb{R} \tag{61}
\end{equation*}
$$

with respect to $\phi$, where $p_{\phi}$ is the projection of the right-hand side of equation (55) along $v_{0}^{\prime}$ (see (57) and (56)). Since the Cahn-Hilliard-Willmore functional is related via Gamma convergence to Willmore's, the principal part of the bifurcation equation turns out to be the linearized Willmore's, appearing in the second variation of the Willmore energy. Recalling (33) from [27], on a hypersurface $\Sigma$ the latter second variation is given by

$$
\mathcal{W}^{\prime \prime}(\Sigma)[\phi, \psi]=\int_{\Sigma}\left(\tilde{L}_{0} \phi\right) \psi d \sigma
$$

where $d \sigma$ is the area form and $\tilde{L}_{0}$ is the self-adjoint operator given by

$$
\begin{aligned}
\tilde{L}_{0} \phi= & L_{0}^{2} \phi+\frac{3}{2} H^{2} L_{0} \phi-H\left(\nabla_{\Sigma} \phi, \nabla_{\Sigma} H\right)+2\left(A \nabla_{\Sigma} \phi, \nabla_{\Sigma} H\right) \\
& +2 H\left\langle A, \nabla^{2} \phi\right\rangle+\phi\left(2\left\langle A, \nabla^{2} H\right\rangle+\left|\nabla_{\Sigma} H\right|^{2}+2 H \operatorname{tr} A^{3}\right) .
\end{aligned}
$$

Here, $L_{0} \phi=-\Delta_{\Sigma} \phi-|A|^{2} \phi$ is the Jacobi operator (related to the second variation of the area functional), $A$ is the second fundamental form, $H$ is the mean curvature and $\operatorname{tr} A^{3}$ is the trace of $A^{3}$. Recalling (6) and (9), on planar curves $\tilde{L}_{0}$ can be written as

$$
\begin{aligned}
\tilde{L}_{0} \phi & =\phi^{(4)}+\frac{5}{2}\left(\phi^{\prime} k^{2}\right)^{\prime}+\left(3-\frac{5}{4} k^{4}\right) \phi \\
& =\phi^{(4)}+\frac{5}{2} k^{2} \phi^{\prime \prime}+5 k k^{\prime} \phi^{\prime}+\left(3-\frac{5}{4} k^{4}\right) \phi
\end{aligned}
$$

Lemma 11. Recalling the definition of the constants (see (37) and (46))

$$
c_{\star}:=\int_{\mathbb{R}}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t>0, \quad d_{\star}:=\int_{\mathbb{R}} t^{2}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t>0
$$

the bifurcation equation can be written in the form

$$
\varepsilon^{-1} c_{\star} \tilde{L}_{0} \phi=c_{\star} \tilde{L}_{0}(h+\psi)=d_{\star} \bar{g}+\varepsilon \mathcal{G}_{\varepsilon}(\psi)
$$

where $\mathcal{G}_{\varepsilon}$ satisfies estimates similar to (36).
Proof. In view of (43) and (38) (see also Section 3.3 for the definition of the $G_{\varepsilon}^{i}$ 's and for the Fourier-truncation $\phi \mapsto \phi_{\star}$ ) one has

$$
\begin{aligned}
& \int_{\mathbb{R}} F\left(\tilde{v}_{\varepsilon, \phi}\right)(s, t) v_{0}^{\prime}(t) \mathrm{d} t=-\varepsilon^{5} c_{\star}\left(\tilde{L}_{0}(h+\psi)+\left(\tilde{L}_{0}-\frac{\mathrm{d}^{4}}{\mathrm{~d} s^{4}}\right)\left(h_{\star}-h+\psi_{\star}-\psi\right)\right) \\
& \quad+\varepsilon^{5} \int_{\mathbb{R}} E_{5}(\varepsilon s, t) v_{0}^{\prime}(t) \mathrm{d} t+2 k \varepsilon^{5}\left(h_{\star}^{\prime}+\psi_{\star}^{\prime}\right)^{2}+\varepsilon^{5} \int_{\mathbb{R}}\left(2 v_{0}^{\prime \prime \prime}-t L_{\star} v_{0}^{\prime \prime}\right) v_{0}^{\prime}(t) \mathrm{d} t \\
& \quad+G_{\varepsilon}^{1}(\psi)+G_{\varepsilon}^{2}(\psi)+G_{\varepsilon}^{3}(\psi)+G_{\varepsilon}^{4}(\psi)
\end{aligned}
$$

where $G_{\varepsilon}^{1}(\psi), G_{\varepsilon}^{2}(\psi), G_{\varepsilon}^{3}(\psi)$ are defined in Section 3.3 and

$$
G_{\varepsilon}^{4}(\psi)(s):=\int_{\mathbb{R}}\left(\chi_{4}(t)-1\right) F\left(\tilde{v}_{\varepsilon, \phi}\right)(s, t) v_{0}^{\prime}(t) \mathrm{d} t
$$

is exponentially small in $\varepsilon$, thus in particular it also satisfies the counterpart of (36). Integrating by parts, it is possible to see that the last term vanishes. By the properties of the smoothing operators (see (26)), the term of order $\varepsilon^{5}$ satisfies

$$
\left\|\left(\tilde{L}_{0}-\frac{\mathrm{d}^{4}}{\mathrm{ds} s^{4}}\right)\left(h_{\star}-h+\psi_{\star}-\psi\right)\right\|_{C^{0, \alpha}(\mathbb{R})} \leqq c \varepsilon^{2}\|h+\psi\|_{C^{4, \alpha}(\mathbb{R})}
$$

since $\tilde{L}_{0}-\frac{\mathrm{d}^{4}}{\mathrm{~d} s^{4}}$ is a second-order differential operator. It remains to deal with the contribution of the term involving $E_{5}$. We compute

$$
\begin{aligned}
c_{\star} & :=\int_{\mathbb{R}}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t=2 \lim _{x \rightarrow \infty} \int_{0}^{x}\left(v_{0}^{\prime}\right)^{2}(t) \mathrm{d} t \\
& =2 \lim _{x \rightarrow \infty} \frac{\left(3 \sinh \left(\frac{x}{\sqrt{2}}\right)+\sinh \left(\frac{3 x}{\sqrt{2}}\right)\right) \operatorname{sech}^{3}\left(\frac{x}{\sqrt{2}}\right)}{6 \sqrt{2}}=\frac{2 \sqrt{2}}{3},
\end{aligned}
$$

and

$$
\begin{aligned}
d_{\star}:= & \int_{\mathbb{R}} t^{2}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t=2 \lim _{x \rightarrow \infty} \int_{0}^{x} t^{2}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t \\
= & 2 \lim _{x \rightarrow \infty} \frac{1}{18}\left(12 \sqrt{2} \operatorname{Li}_{2}\left(-\mathrm{e}^{-\sqrt{2} x}\right)-6 \sqrt{2} x^{2}+6 \sqrt{2} x^{2} \tanh \left(\frac{x}{\sqrt{2}}\right)\right. \\
& -24 x \log \left(\mathrm{e}^{-\sqrt{2} x}+1\right)-6 \sqrt{2} \tanh \left(\frac{x}{\sqrt{2}}\right) \\
& \left.+3 x\left(\sqrt{2} x \tanh \left(\frac{x}{\sqrt{2}}\right)+2\right) \operatorname{sech}^{2}\left(\frac{x}{\sqrt{2}}\right)+\sqrt{2} \pi^{2}\right)=\frac{\sqrt{2}}{9}\left(\pi^{2}-6\right)
\end{aligned}
$$

Here $\mathrm{Li}_{2}$ is the dilogarithmic function, defined as

$$
\operatorname{Li}_{2}(x)=-\int_{1}^{x} \frac{\log t}{t-1} \mathrm{~d} t
$$

and it satisfies

$$
\begin{aligned}
& \frac{\mathrm{d}}{\mathrm{~d} x} \mathrm{Li}_{2}\left(-\mathrm{e}^{-\sqrt{2} x}\right)=\sqrt{2} \log \left(\mathrm{e}^{-\sqrt{2} x}+1\right), \quad \text { with } \\
& \mathrm{Li}_{2}(0)=0, \mathrm{Li}_{2}(-1)=-\frac{\pi^{2}}{12}
\end{aligned}
$$

For these and further details about the dilogarithmic function, see for instance [1, p. 1004].

Now we deal with the projection of $E_{5}$. Integrating by parts, it is possible to see that

$$
\begin{aligned}
\int_{\mathbb{R}} t^{3} v_{0}^{\prime \prime} v_{0}^{\prime} \mathrm{d} t & =-\frac{3}{2} d_{\star} \\
\int_{\mathbb{R}} \eta^{\prime} v_{0}^{\prime} \mathrm{d} t & =\frac{1}{4} d_{\star} \\
2 \int_{\mathbb{R}}\left(L_{\star} \eta_{3}\right)^{\prime} v_{0}^{\prime} \mathrm{d} t & =-2 \int_{\mathbb{R}} L_{\star} \eta_{3} v_{0}^{\prime \prime} \mathrm{d} t=\int_{\mathbb{R}} L_{\star} \eta_{3} L_{\star}\left(t v_{0}^{\prime}\right) \mathrm{d} t \\
& =\int_{\mathbb{R}} L_{\star}^{2} \eta_{3} t v_{0}^{\prime} \mathrm{d} t=d_{\star}
\end{aligned}
$$

and that

$$
k\left(k^{\prime}\right)^{2} \int_{\mathbb{R}}\left(2 t\left(L_{\star} \eta_{1}\right)^{\prime}-9 t^{2} v_{0}^{\prime}-4 \eta^{\prime}-18 L_{\star} \eta_{1}\right) v_{0}^{\prime} \mathrm{d} t=-9 k\left(k^{\prime}\right)^{2} d_{\star}
$$

Moreover, since, due to (41),

$$
\int_{\mathbb{R}} Q(\eta, \eta) t v_{0}^{\prime} \mathrm{d} t=-12 \int_{\mathbb{R}} v_{0} v_{0}^{\prime \prime} \eta^{2} \mathrm{~d} t+6 \int_{\mathbb{R}} t^{2}\left(v_{0}^{\prime}\right)^{2} v_{0} \eta \mathrm{~d} t
$$

we have

$$
\begin{aligned}
& 2 \int_{\mathbb{R}}\left(L_{\star} \eta_{2}\right)^{\prime} v_{0}^{\prime}=\int_{\mathbb{R}} L_{\star}^{2} \eta_{2} t v_{0}^{\prime} \mathrm{d} t \\
& =-\frac{3}{2} d_{\star}-\int_{\mathbb{R}} t \eta^{\prime \prime} v_{0}^{\prime} \mathrm{d} t+6 \int_{\mathbb{R}} v_{0} v_{0}^{\prime \prime} \eta^{2} \mathrm{~d} t-3 \int_{\mathbb{R}} t^{2}\left(v_{0}^{\prime}\right)^{2} \eta v_{0} \mathrm{~d} t .
\end{aligned}
$$

The quadratic term in $\eta$ gives

$$
12 \int_{\mathbb{R}} \eta \eta^{\prime} v_{0}^{\prime} v_{0} \mathrm{~d} t+6 \int_{\mathbb{R}} \eta^{2}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t=-6 \int_{\mathbb{R}} v_{0} v_{0}^{\prime \prime} \eta^{2} \mathrm{~d} t
$$

With similar reasoning, the quadratic term containing $\eta$ and $\eta_{1}$ gives

$$
\begin{aligned}
\int_{\mathbb{R}} Q\left(\eta, \eta_{1}\right) v_{0}^{\prime} \mathrm{d} t= & -3 \int_{\mathbb{R}} v_{0}\left(v_{0}^{\prime}\right)^{2} \eta t^{2} \mathrm{~d} t+\frac{3 d_{\star}}{c_{\star}} \int_{\mathbb{R}} v_{0}\left(v_{0}^{\prime}\right)^{2} \eta \mathrm{~d} t \\
& +3 \int_{\mathbb{R}} t v_{0}\left(v_{0}^{\prime}\right)^{2} \eta_{1} \mathrm{~d} t
\end{aligned}
$$

Moreover,

$$
\int_{\mathbb{R}} t\left(L_{\star} \eta_{1}\right)^{\prime} v_{0}^{\prime} \mathrm{d} t=-\frac{d_{\star}}{2}
$$

We note that

$$
\begin{aligned}
& L_{\star}\left(t v_{0} v_{0}^{\prime} / \sqrt{2}\right)=v_{0}^{\prime \prime \prime}+3 t v_{0}\left(v_{0}^{\prime}\right)^{2} \\
& L_{\star}\left(t v_{0}^{\prime}\left(1+\sqrt{2} t v_{0}\right) / 4\right)=t v_{0}^{\prime \prime \prime}+\frac{3}{2} t^{2} v_{0}\left(v_{0}^{\prime}\right)^{2} \\
& L_{\star}\left(v_{0} v_{0}^{\prime} / 3 \sqrt{2}\right)=v_{0}\left(v_{0}^{\prime}\right)^{2}
\end{aligned}
$$

thus

$$
\begin{align*}
\int_{\mathbb{R}} \eta_{1}\left(v_{0}^{\prime \prime \prime}+3 t v_{0}\left(v_{0}^{\prime}\right)^{2}\right) \mathrm{d} t & =\frac{1}{\sqrt{2}} \int_{\mathbb{R}} t v_{0} v_{0}^{\prime} L_{\star} \eta_{1} \mathrm{~d} t=-\frac{d_{\star}}{4},  \tag{62}\\
\int_{\mathbb{R}} \eta\left(t v_{0}^{\prime \prime \prime}+\frac{3}{2} t^{2} v_{0}\left(v_{0}^{\prime}\right)^{2}\right) \mathrm{d} t & =\frac{1}{4} \int_{\mathbb{R}} t v_{0}^{\prime}\left(1+\sqrt{2} t v_{0}\right) L_{\star} \eta \mathrm{d} t=\frac{5}{16} d_{\star},  \tag{63}\\
\int_{\mathbb{R}} \eta v_{0}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t & =\frac{1}{3 \sqrt{2}} \int_{\mathbb{R}} v_{0} v_{0}^{\prime} L_{\star} \eta \mathrm{d} t=\frac{1}{18 \sqrt{2}} . \tag{64}
\end{align*}
$$

In order to prove (62), (63) and (64) we observe that, concerning the first integral,

$$
\begin{aligned}
& \frac{1}{\sqrt{2}} \int_{0}^{x} t v_{0} v_{0}^{\prime} L_{\star} \eta_{1} \mathrm{~d} t=\frac{1}{288}\left(-72 \sqrt{2} \mathrm{Li}_{2}\left(-\mathrm{e}^{-\sqrt{2} x}\right)-6 \sqrt{2}\left(\pi^{2}-6 x^{2}\right)\right. \\
& \quad+2 \sqrt{2}\left(-18 x^{2}+\pi^{2}+12\right) \tanh \left(\frac{x}{\sqrt{2}}\right)
\end{aligned}
$$

$$
\begin{align*}
& +3 x\left(6 x^{2}-\pi^{2}+6\right) \operatorname{sech}^{4}\left(\frac{x}{\sqrt{2}}\right) \\
& +\left(\sqrt{2}\left(-18 x^{2}+\pi^{2}-6\right) \tanh \left(\frac{x}{\sqrt{2}}\right)-36 x\right) \operatorname{sech}^{2}\left(\frac{x}{\sqrt{2}}\right) \\
& \left.+144 x \log \left(\mathrm{e}^{-\sqrt{2} x}+1\right)\right) \tag{65}
\end{align*}
$$

Concerning the second integral, one has indeed that

$$
\begin{align*}
& \frac{1}{4} \int_{0}^{x} t v_{0}^{\prime}\left(1+\sqrt{2} t v_{0}\right) L_{\star} \eta \mathrm{d} t=\frac{1}{288}\left(60 \sqrt{2} \mathrm{Li}_{2}\left(-\mathrm{e}^{-\sqrt{2} x}\right)-9 x^{3} \operatorname{sech}^{4}\left(\frac{x}{\sqrt{2}}\right)\right. \\
& \quad+5\left(\sqrt{2}\left(\pi^{2}-6 x^{2}\right)+6 \sqrt{2}\left(x^{2}-1\right) \tanh \left(\frac{x}{\sqrt{2}}\right)\right. \\
& \left.\quad-24 x \log \left(\mathrm{e}^{-\sqrt{2} x}+1\right)\right) \\
& \left.\quad+15 x\left(\sqrt{2} x \tanh \left(\frac{x}{\sqrt{2}}\right)+2\right) \operatorname{sech}^{2}\left(\frac{x}{\sqrt{2}}\right)\right) \tag{66}
\end{align*}
$$

For the third integral, one has that
$\frac{1}{3 \sqrt{2}} \int_{0}^{x} v_{0} v_{0}^{\prime} L_{\star} \eta \mathrm{d} t=\frac{(-6 \sqrt{2} x+4 \sinh (\sqrt{2} x)+\sinh (2 \sqrt{2} x)) \operatorname{sech}^{4}\left(\frac{x}{\sqrt{2}}\right)}{288 \sqrt{2}}$.
Taking the sum, we get

$$
\begin{equation*}
\int_{\mathbb{R}} E_{5}(s, t) v_{0}^{\prime}(t) \mathrm{d} t=\frac{9}{8} d_{\star} k^{5}(s)-9 k\left(k^{\prime}\right)^{2}(s) d_{\star}=d_{\star} \bar{g} . \tag{67}
\end{equation*}
$$

To conclude the proof we observe that, thanks to Propositions 9 and 10,

$$
G_{\varepsilon}^{5}(\psi)(s):=\varepsilon^{-6} \int_{\mathbb{R}} \mathrm{T}\left(U_{\phi}, V_{\varepsilon, \phi, U}, \phi\right)\left(\varepsilon^{-1} s, t\right) v_{0}^{\prime}(t) \mathrm{d} t
$$

satisfies (36). Therefore, if we set

$$
\begin{align*}
\mathcal{G}_{\varepsilon}(\psi):= & \varepsilon^{-2}\left(\frac{\mathrm{~d}^{4}}{\mathrm{~d} s^{4}}-\tilde{L}_{0}\right)\left(h_{\star}-h+\psi_{\star}-\psi\right)+G_{\varepsilon}^{1}(\psi)+G_{\varepsilon}^{2}(\psi)+G_{\varepsilon}^{3}(\psi) \\
& +G_{\varepsilon}^{4}(\psi)+G_{\varepsilon}^{5}(\psi) \tag{68}
\end{align*}
$$

the Lemma is proved.
Proposition 12. For $\varepsilon>0$ small enough, the bifurcation equation (61) admits a solution $\phi \in C_{\bar{T}}^{4, \alpha}(\mathbb{R})$, such that

$$
\phi=\varepsilon \frac{d_{\star}}{c_{\star}} \bar{\phi}+\varepsilon \psi
$$

with $\psi \in C_{\bar{T}}^{4, \alpha}$ fulfilling $\|\psi\|_{C^{4, \alpha}(\mathbb{R})} \leqq c \varepsilon$ for some constant $c>0$ and where $c_{\star}, d_{\star}$ are given by (37) and (46).

Proof. We recall that we look for a solution of the form $\phi=\varepsilon(h+\psi)$ and, by Lemma 11, the bifurcation equation can be written in the form

$$
\tilde{L}_{0}(h+\psi)=\bar{g}+\varepsilon \mathcal{G}_{\varepsilon}(\psi),
$$

where $\mathcal{G}_{\varepsilon}$ is given by (68). In order to solve this, first we set $h:=\frac{d_{\star}}{c_{\star}} \bar{\phi}$ in such a way that

$$
\tilde{L}_{0} h=\frac{d_{\star}}{c_{\star}} \bar{g}
$$

(see Proposition 4), then we treat the fixed point problem

$$
\tilde{L}_{0} \psi=\varepsilon \mathcal{G}_{\varepsilon}(\psi)
$$

using the inverse of $\tilde{L}_{0}$ constructed in Proposition 3. In order to apply the contraction mapping theorem, we need to prove the Lipschitz character of $\mathcal{G}_{\varepsilon}$. This follows from the definitions of $p_{\phi}$ and T (see (57) and (56)), the Lipschitz regularity of $G_{\varepsilon}^{i}, i=1, \ldots, 5$ (they all meet (36)), which follows from property (35), satisfied by $F_{\varepsilon}^{1}, F_{\varepsilon}^{2}$ and $F_{\varepsilon}^{3}$ and the Lipschitz dependence of $U$ and $V$ on the datum $\phi$ (see Propositions 10 and 9).

### 4.3. Proof of Theorem 1

Thanks to the results in the previous subsections the proving existence of a symmetric solution to (1), we only need to prove (5). By the symmetries of $u_{T}$, we can reduce ourselves to study the sign of $\partial_{x_{2}} u$ in the strip $\left\{0 \leqq x_{1} \leqq\left(\gamma_{T}\right)_{1}(T / 4)\right\}$.

Before proceeding, similarly to (17), for any $x \in V_{\varepsilon}$ (see (18)) we set

$$
x=\tilde{Z}_{\varepsilon}(s, z)=\gamma_{T}(s)+z \gamma_{T}^{\prime}(s)^{\perp} ; \quad \gamma_{T}(s)=\frac{1}{\varepsilon} \gamma(\varepsilon s) .
$$

Since $\gamma_{T}(s)=\gamma^{\prime}(\varepsilon s)$, the latter formula becomes

$$
\tilde{Z}_{\varepsilon}(s, z)=\frac{1}{\varepsilon} \gamma(\varepsilon s)+z \gamma^{\prime}(\varepsilon s)^{\perp}
$$

We would like to understand the inverse function, namely the dependence of $(t, z)$ on ( $x_{1}, x_{2}$ ), especially near the $x_{1}$-axis. We notice first that $\tilde{Z}_{\varepsilon}(0, z)=(z, 0)$, and that

$$
\left(\begin{array}{cc}
\partial_{s} x_{1} & \partial_{s} x_{2} \\
\partial_{z} x_{1} & \partial_{z} x_{2}
\end{array}\right)=\left(\begin{array}{cc}
-(1-\varepsilon z k(\varepsilon s)) \sin \int_{0}^{\varepsilon s} k(\tau) d \tau & (1-\varepsilon z k(\varepsilon s)) \cos \int_{0}^{\varepsilon s} k(\tau) d \tau \\
\cos \int_{0}^{\varepsilon_{0}^{s s}} k(\tau) d \tau & \sin \int_{0}^{\varepsilon s} k(\tau) d \tau
\end{array}\right) .
$$

Recalling that $k(0)=0$ and that $k^{\prime}(0)<0$, differentiating the definition of $\tilde{Z}_{\varepsilon}$ and taking the scalar product with $\gamma^{\prime}(\varepsilon s)^{\perp}$, it is easy to see that $\partial_{x_{2}} z=\gamma_{1}^{\prime}(\varepsilon s)$ in $V_{\varepsilon}$, then near the origin one has, for $\bar{\delta}>0$ small

$$
\begin{equation*}
\partial_{x_{2}} z=\frac{\left|k^{\prime}(0)\right|}{2} x_{2}^{2} \varepsilon^{2}\left(1+o_{\varepsilon}(1)\right) ; \quad x \in V_{\varepsilon}, \quad\left|x_{2}\right|<\frac{\bar{\delta}}{\varepsilon} \tag{69}
\end{equation*}
$$

After these preliminaries, we have

Proposition 13. Let $\tilde{v}_{\varepsilon, \phi}(s, t)$ be the approximate solution defined in (42). Then there exists a fixed constant $C$ such that

$$
\frac{\partial \tilde{v}_{\varepsilon, \phi}}{\partial x_{2}} \geqq-C \varepsilon^{3} \quad \text { in } V_{\varepsilon}
$$

Proof. Recall that in $V_{\varepsilon}$ we defined

$$
\tilde{v}_{\varepsilon, \phi}(s, t):=v_{0}(t)+v_{1, \varepsilon, \phi}(s, t)+v_{2, \varepsilon, \phi}(s, t)+v_{3, \varepsilon, \phi}(s, t)+v_{4, \varepsilon, \phi}(s, t) .
$$

We begin by estimating the $x_{2}$-derivative of the first term. Recalling that $t=$ $z-\phi_{\star}(\varepsilon s)$, we have

$$
\partial_{x_{2}} v_{0}(t)=v_{0}^{\prime}(t)\left[\partial_{x_{2}} z-\varepsilon \phi_{\star}^{\prime}(\varepsilon s) \partial_{x_{2}} s\right]
$$

Concerning the function $\phi_{\star}^{\prime}$ we recall that by (26), for $\alpha \in(0,1)$ and $\theta=\frac{1}{\varepsilon}$ one has

$$
\left\|\phi-\phi_{\star}\right\|_{C^{1, \alpha}} \leqq C \varepsilon^{3}\|\phi\|_{C^{4, \alpha}} \leqq C \varepsilon^{4}
$$

Moreover, by Proposition 12 we had that

$$
\left\|\phi-\varepsilon \frac{d_{\star}}{c_{\star}} \bar{\phi}\right\|_{C^{4, \alpha}} \leqq C \varepsilon^{2}
$$

The latter two formulas imply that near the origin

$$
\varepsilon \phi_{\star}^{\prime}(\varepsilon s) \geqq-C \varepsilon^{3}
$$

and therefore also that near the $x_{1}$ axis, by (69),

$$
\partial_{x_{2}} v_{0}(t) \geqq-C \varepsilon^{3}+\frac{1}{2} x_{2}^{2} \varepsilon^{2}\left(1+o_{\varepsilon}(1)\right)
$$

Concerning instead $v_{1, \varepsilon, \phi}$, defined in (27), we have that

$$
\begin{aligned}
& \partial_{x_{2}} v_{1, \varepsilon, \phi}=\left(v_{0}^{\prime}\left(t+\phi_{\star}(\varepsilon s)-\phi(\varepsilon s)\right)-v^{\prime}(t)\right)\left[\partial_{x_{2}} z-\varepsilon \phi_{\star}^{\prime}(\varepsilon s) \partial_{x_{2}} s\right] \\
&+\varepsilon v_{0}^{\prime}\left(t+\phi_{\star}-\phi\right)\left(\phi_{\star}^{\prime}(\varepsilon s)-\phi^{\prime}(\varepsilon s)\right) \partial_{x_{2}} s .
\end{aligned}
$$

This term can be estimated by

$$
C\left\|\phi_{\star}-\phi\right\|_{L^{\infty}\left|\partial_{x_{2}} z-\varepsilon \phi_{\star}^{\prime}(\varepsilon s) \partial_{x_{2}} s\right|+C \varepsilon\left\|\phi_{\star}-\phi\right\|_{C^{1, \alpha}} . . . . ~}
$$

Using (26) we can check that this term is of order $\varepsilon^{5}$.
We turn next to $v_{2, \varepsilon, \phi}$, see (28). The first summand in its definition is quite easy to treat. The terms $\varepsilon^{3} L \phi_{\star}(\varepsilon s) \eta(t)$ and $\varepsilon^{2} \phi_{\star}^{\prime}(\varepsilon s)^{2} \tilde{\eta}(t)$, involving the Fourier truncation $\phi_{\star}$ might seem more delicate. However, being $L$ of second order (see (29)), using (26) and recalling that $\|\phi\|_{C^{4, \alpha}} \leqq c \varepsilon$, one has that the $x_{2}$-derivative of both these terms is of order $\varepsilon^{4}$.

All other terms in $\tilde{v}_{\varepsilon, \phi}$ can be estimated quite easily, and it is also straightforward to show the monotonicity of $\tilde{v}_{\varepsilon, \phi}$ in $x_{2}$ in $V_{\varepsilon}$ for $\left|x_{2}\right| \geqq \frac{\bar{\delta}}{\varepsilon}$, since in this region $v_{0}(t)$ has $x_{2}$-derivative bounded away from zero.

Proof of Theorem 1 completed. We notice that the solution $u_{T}$ is obtained by multiplying $\tilde{v}_{\varepsilon, \phi}$ by a cut-off function (not identically equal to 1 in a region where $\tilde{v}_{\varepsilon, \phi}$ is exponentially small in $\varepsilon$ ) and by adding a correction $w$ which is of order $\varepsilon^{5}$ in $C^{1}$ norm, see the beginning of Section 4. Then (5) follows from Proposition 13.

The weighted norm estimate on the correction $w$, the fact that $v_{0}^{\prime}(t)$ has non zero gradient for $t$ close to zero, and the construction of $v_{\varepsilon, \phi}$ also imply (4) by a direct application of the implicit function theorem.

## 5. Proof of Some Technical Results

In this section we collect the proofs of some technical results, most notably of Propositions 9 and 10.

### 5.1. Proof of Proposition 9

Our main strategy is the following: if $\Gamma_{\varepsilon, \phi}$ is as in Remark 8, we first we study the linear equation

$$
\begin{equation*}
\left(-\Delta+\Gamma_{\varepsilon, \phi}\right)^{2} u=f \text { in } \mathbb{R}^{2} \tag{70}
\end{equation*}
$$

where $f$ is a fixed function with finite $C_{L, \delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)$ norm (see (34)), that is decaying away from the curve $\gamma_{T}$ at rate $\mathrm{e}^{-\delta d\left(\cdot, \gamma_{T}\right)}$, even and periodic with period $L$ in $x_{1}$ (see Section 4.2). The aim is to construct a right inverse of the operator $\left(-\Delta+\Gamma_{\varepsilon, \phi}\right)^{2}$, in order to solve equation (54) by a fixed point argument (see Section 4.2).

In order to treat equation (54), we will endow the space $C_{L, \delta}^{n, \alpha}\left(\mathbb{R}^{2}\right)$ with the norm introduced in (33).
5.1.1. The Linear Problem In order to solve (70), we first consider the second order equation

$$
\begin{equation*}
-\Delta u+\Gamma_{\varepsilon, \phi} u=f \text { in } \mathbb{R}^{2} \tag{71}
\end{equation*}
$$

proving the following result (recall (13)):
Lemma 14. Let $f \in C_{L, \delta}^{0, \alpha}\left(\mathbb{R}^{2}\right), 0<\delta<\sqrt{2}$. Then, for $\varepsilon$ small enough and $\phi \in C_{T}^{4, \alpha}(\mathbb{R})$ with $\|\phi\|_{C^{4, \alpha}(\mathbb{R})}<1$, equation (71) admits a unique solution $u:=$ $\tilde{\Psi}_{\varepsilon, \phi}(f) \in C_{L, \delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)$ satisfying $\|u\|_{C_{\delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c\|f\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}$, for some constant $c>0$ independent of $\varepsilon$ and $\phi$.

Proof. Step (i): existence, uniqueness and local Hölder regularity on a strip.
Recalling that $L$ is the $x_{1}$-period of $\gamma_{T}$, define the strip

$$
S=(-L / 2, L / 2) \times \mathbb{R}
$$

First we look for a solution to the Neumann problem

$$
\begin{cases}-\Delta u+\Gamma_{\varepsilon, \phi} u=f & \text { in } S  \tag{72}\\ \partial_{x_{1}} u=0 & \text { on } \partial S\end{cases}
$$

and then we will extend it by periodicity to the whole $\mathbb{R}^{2}$. By definition, $w \in H^{1}(S)$ is a weak solution to problem (72) if

$$
\begin{equation*}
\int_{S}\langle\nabla w, \nabla v\rangle \mathrm{d} x+\int_{S} \Gamma_{\varepsilon, \phi} w v \mathrm{~d} x=\int_{S} f v \mathrm{~d} x, \quad \text { for any } v \in H^{1}(S) \tag{73}
\end{equation*}
$$

The existence and uniqueness of such a solution follows from the Riesz representation theorem. Since $f \in C^{0, \alpha}(\bar{S})$, it follows that $w \in C^{2, \alpha}(\bar{S})$. Moreover, choosing an arbitrary test function $v \in C^{1}(\bar{S})$ and applying the divergence theorem, we can see that

$$
\int_{S} v\left(-\Delta w+\Gamma_{\varepsilon, \phi} w\right) \mathrm{d} x+\int_{\partial S} \partial_{x_{1}} u v \mathrm{~d} x_{1}=\int_{S} f v \mathrm{~d} x .
$$

Taking $v \in C_{0}^{\infty}(S)$, the PDE is satisfied in the classical sense. Taking once again $v \in C^{1}(\bar{S})$, we also have that $\partial_{x_{1}} w=0$ on $\partial S$.

Step (ii): Symmetry and extension to an entire solution
By the symmetries of the Laplacian and the uniqueness of the solution, if $f$ is even in $x_{1}$ then the same is true for $w$, thus $w\left(-L / 2, x_{2}\right)=w\left(L / 2, x_{2}\right)$ and $\partial_{x_{1}}^{2} w\left(-L / 2, x_{2}\right)=\partial_{x_{1}}^{2} w\left(L / 2, x_{2}\right)$. As a consequence, it is possible to extend $w$ by periodicity to an entire solution $u \in C^{2, \alpha}\left(\mathbb{R}^{2}\right)$.

Step (iii): $u \in L^{\infty}\left(\mathbb{R}^{2}\right)$.
By elliptic estimates and the Sobolev embeddings,

$$
\begin{aligned}
\|u\|_{L^{\infty}\left(B_{1}(x)\right)} & \leq c\|u\|_{W^{2,2}\left(B_{1}(x)\right)} \leqq c\left(\|u\|_{L^{2}\left(B_{2}(x)\right)}+\|f\|_{L^{2}\left(B_{2}(x)\right)}\right) \\
& \leq c\left(\|w\|_{L^{2}(S)}+\|f\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}\right)<\infty
\end{aligned}
$$

for any $x \in \mathbb{R}^{2}$, thus $u \in L^{\infty}\left(\mathbb{R}^{2}\right)$.
Step (iv): Decay of the solution: $u \varphi_{\varepsilon, \delta}^{-1} \in L^{\infty}\left(\mathbb{R}^{2}\right)$ (see (32) for the definition of $\left.\varphi_{\varepsilon, \delta}\right), 0<\delta<\sqrt{2}$.

For suitable constants $\lambda, \tau>0$, we will use the function $\lambda \varphi+\tau \varphi^{-1}$ as a barrier, where we have set $\varphi:=\varphi_{\varepsilon, \delta}$. More precisely, we fix $\rho>0$ and $z \in \mathbb{R}^{2}$ with $d\left(z, \gamma_{T}\right)>\rho$. Then we fix $\tau>0$ small and $R>\left|d\left(z, \gamma_{T}\right)\right|$. Therefore $u$ fulfils

$$
u(x)<\|u\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}<\lambda \varphi(x)<\lambda \varphi(x)+\tau \varphi^{-1}(x)
$$

if $d\left(x, \gamma_{T}\right)=\rho$, provided $\lambda>\|u\|_{L^{\infty}\left(\mathbb{R}^{2}\right)} \sup _{d\left(x, \gamma_{\varepsilon}\right)=\rho} \varphi^{-1}>0$. Furthermore,

$$
u<\|u\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}<\tau \varphi^{-1}<\lambda \varphi+\tau \varphi^{-1}
$$

if $d\left(x, \gamma_{T}\right)=R$, provided $R$ is large enough. Moreover,

$$
\begin{gathered}
\left(-\Delta+\Gamma_{\varepsilon, \phi}\right)\left(u-\left(\lambda \varphi+\tau \varphi^{-1}\right)\right) \leq\left(c-\lambda\left(\Gamma_{\varepsilon, \phi}-\delta^{2}\right)\right) \varphi \\
-\tau \varphi^{-1}\left\{\Gamma_{\varepsilon, \phi}+\delta^{2}-2 \frac{|\nabla \varphi|^{2}}{\varphi^{2}}\right\}<0 \text { for } x \in \Omega,
\end{gathered}
$$

where $\Omega:=\left\{x: \rho<d\left(x, \gamma_{T}\right)<R\right\}$, if $\lambda$ is large enough. We observe that, if we fix $0<\beta<W^{\prime \prime}(1)-\delta^{2}$, then, for $\varepsilon$ small enough,

$$
\begin{align*}
\Gamma_{\varepsilon, \phi}-\delta^{2} & =\left(\chi_{1}-1\right)\left(W^{\prime \prime}(1)-W^{\prime \prime}\left(\tilde{v}_{\varepsilon, \phi}\right)\right)+W^{\prime \prime}(1)-\delta^{2} \\
& >W^{\prime \prime}(1)-\delta^{2}-\beta>0 \tag{74}
\end{align*}
$$

Thus the function $c /\left(\Gamma_{\varepsilon, \phi}-\delta^{2}\right)$ is bounded from above, therefore we can take $\lambda>\sup _{x \in \Omega} c /\left(\Gamma_{\varepsilon, \phi}-\delta^{2}\right)$. The term multiplying $-\tau \varphi^{-1}$ is positive, due to the estimate $|\nabla \varphi|^{2} / \varphi^{2} \leqq \delta^{2}$ and (74). Therefore, by the maximum principle we get that $u(z)<\lambda \varphi+\tau \varphi^{-1}$, in the complement of the region $\{|t| \leq \rho\}$ and for any $\tau>0$. In the same way, one can prove that $u(z)>-\lambda \varphi-\tau \varphi^{-1}$. Letting $\sigma \rightarrow 0$, we get that $u \varphi^{-1} \in L^{\infty}\left(\mathbb{R}^{2}\right)$.

Step (v): estimate of the $L^{\infty}$-norm of $и \varphi_{\varepsilon, \delta}^{-1}$.
Let us set $\tilde{u}:=u \varphi_{\varepsilon, \delta}^{-1}$. It is possible to show that

$$
\begin{align*}
\left(-\Delta+\Gamma_{\varepsilon, \phi}\right) \tilde{u} & =\tilde{f}-2\left(\nabla u, \nabla \varphi^{-1}\right)-u \Delta \varphi^{-1} \\
& =\tilde{f}-2 \varphi\left(\nabla \tilde{u}, \nabla \varphi^{-1}\right)+\tilde{u}\left(2 \frac{|\nabla \varphi|^{2}}{\varphi^{2}}-\varphi \Delta \varphi^{-1}\right) \\
& =\tilde{f}-2 \varphi\left(\nabla \tilde{u}, \nabla \varphi^{-1}\right)+\delta^{2} \tilde{u} \tag{75}
\end{align*}
$$

(once again, we have set $\varphi:=\varphi_{\varepsilon, \delta}$ in the above computation). First we assume that there exists a point $y \in \mathbb{R}^{2}$ such that $|\tilde{u}(y)|=\max _{x \in \mathbb{R}^{2}}|\tilde{u}(x)|$. If $\tilde{u}(y)>0$, then $y$ is a maximum point, thus $\nabla \tilde{u}(y)=0$ and

$$
\left(\Gamma_{\varepsilon, \phi}(y)-\delta^{2}\right) \tilde{u}(y) \leq-\Delta \tilde{u}(y)+\left(\Gamma_{\varepsilon, \phi}(y)-\delta^{2}\right) \tilde{u}(y)=\tilde{f}(y),
$$

and therefore

$$
\|\tilde{u}\|_{L^{\infty}\left(\mathbb{R}^{2}\right)} \leqq c\|\tilde{f}\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}
$$

A similar argument shows that the same estimate is true if $\tilde{u}(y)<0$ (a minimum point).

If the maximum is not achieved, then there exists a sequence $\left(x_{k}\right)_{k} \subset \mathbb{R}^{2}$ such that $\left|\tilde{u}\left(x_{k}\right)\right| \rightarrow \sup _{x \in \mathbb{R}^{2}}|\tilde{u}(x)|$. Then we define $\tilde{u}_{k}(x):=\tilde{u}\left(x+x_{k}\right)$. Up to a subsequence, $\tilde{u}_{k} \rightarrow w$ in $C_{l o c}^{2}\left(\mathbb{R}^{2}\right)$, and $w$ is still a solution to the linear equation

$$
-\Delta w+\Gamma_{\varepsilon, \phi} w=\tilde{f} \text { in } \mathbb{R}^{2}
$$

Moreover, $|w(0)|=\|w\|_{L^{\infty}\left(\mathbb{R}^{2}\right)} \leq\|\tilde{f}\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}$. As a consequence, by the previous step,

$$
\|\tilde{u}\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}=\|w\|_{L^{\infty}\left(\mathbb{R}^{2}\right)} \leq\|\tilde{f}\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}
$$

Step (vi): Decay of the derivatives.
By (75), step (v) and [25] (Chapter 6.1, Corollary 6.3),

$$
\|\tilde{u}\|_{C^{2, \alpha}\left(B_{1}(x)\right)} \leqq c\left(\|\tilde{u}\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}+\|\tilde{f}\|_{C^{0, \alpha}\left(\mathbb{R}^{2}\right)}\right) \leqq c\|\tilde{f}\|_{C^{0, \alpha}\left(\mathbb{R}^{2}\right)}<\infty
$$

for any $x \in \mathbb{R}^{2}$, thus $u \in C_{\sigma}^{2, \alpha}\left(\mathbb{R}^{2}\right)$ and

$$
\|\tilde{u}\|_{C^{2, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c\|\tilde{f}\|_{C^{0, \alpha}\left(\mathbb{R}^{2}\right)}
$$

This concludes the proof.

Lemma 15. Let $f \in C_{L, \delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)$, with $0<\delta<\sqrt{2}$. Then, for $\varepsilon$ small enough and $\phi \in C_{T}^{4, \alpha}(\mathbb{R})$ with $\|\phi\|_{C^{4, \alpha}(\mathbb{R})}<1$, equation (70) admits a unique solution $V:=\Psi_{\varepsilon, \phi}(f) \in C_{L, \delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)$ satisfying the estimate $\|V\|_{C_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c\|f\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}$ for some constant $c>0$ independent of $\varepsilon$ and $\phi$.

Proof. Given $f \in C_{L, \delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)$, we have to find $V \in C_{L, \delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)$ fulfilling

$$
\left\{\begin{array}{l}
\left(-\Delta+\Gamma_{\varepsilon, \phi}\right)^{2} V=f \\
\|V\|_{C_{\delta}^{4, \alpha}\left(\mathbb{R}^{3}\right)} \leqq c\|f\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}
\end{array}\right.
$$

In order to do so, we use Lemma 14 twice to find $u \in C_{L, \delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)$ and $V \in C_{L, \delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)$, such that

$$
\left\{\begin{array}{l}
\left(-\Delta+\Gamma_{\varepsilon, \phi}\right) u=f \\
\left(-\Delta+\Gamma_{\varepsilon, \phi}\right) V=u
\end{array}\right.
$$

and

$$
\left\{\begin{array}{l}
\|u\|_{C_{\delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c\|f\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \\
\|V\|_{C_{\delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c\|u\|_{\delta}^{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} .
\end{array}\right.
$$

Now it remains to estimate the higher-order derivatives of $u$. For this purpose, we differentiate the equation satisfied by $u$ to get

$$
\left(-\Delta+\Gamma_{\varepsilon, \phi}\right) V_{j}=u_{j}-\left(\Gamma_{\varepsilon, \phi}\right)_{j} V
$$

for $j=1, \ldots, 3$. By Proposition 14 , we can find a unique solution $\tilde{V} \in C_{L, \delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)$ such that

$$
\|\tilde{V}\|_{C_{\delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c\left(\left\|u_{j}\right\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}+\|f\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}\right) \leqq c\|f\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)},
$$

hence $\tilde{V}=V_{j}$.
Similarly, differentiating the equation once again, we see that

$$
\left(-\Delta+\Gamma_{\varepsilon, \phi}\right) V_{i j}=u_{i j}-\left(\Gamma_{\varepsilon, \phi}\right)_{i} V_{j}-\left(\Gamma_{\varepsilon, \phi}\right)_{j} V_{i}-\left(\Gamma_{\varepsilon, \phi}\right)_{i j} V
$$

for $i, j=1, \ldots, 3$, so in particular $V_{i j} \in C_{L, \delta}^{4, \alpha}\left(\mathbb{R}^{3}\right)$ and

$$
\left\|V_{i j}\right\|_{C_{\delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c\left(\left\|u_{i j}\right\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}+\|f\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}\right) \leqq c\|f\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}
$$

This concludes the proof.
5.1.2. A Fixed Point Argument Equation (54), whose resolvability is the purpose of this subsection, is equivalent to the fixed point problem

$$
\begin{aligned}
V=\mathcal{S}_{1}(V):=- & \Psi_{\varepsilon, \phi}\left\{\left(1-\chi_{2}\right) F\left(v_{\varepsilon, \phi}\right)+\left(1-\chi_{1}\right) Q_{\varepsilon, \phi}\left(\chi_{2} U+V\right)\right. \\
& \left.+\mathrm{N}_{\varepsilon, \phi}(U)+\mathrm{P}_{\varepsilon, \phi}(V)\right\}
\end{aligned}
$$

We will solve by showing that $\mathcal{S}_{1}$ is a contraction on the ball

$$
\Lambda_{1}:=\left\{V \in C_{L, \delta}^{4, \alpha}\left(\mathbb{R}^{3}\right):\|V\|_{C_{\delta}^{4, \alpha}\left(\mathbb{R}^{3}\right)} \leqq C_{1} \mathrm{e}^{-\delta / 8 \varepsilon}\right\}
$$

provided the constant $C_{1}$ is large enough. This step of the proof is similar to that in Section 6.2 of [35]. In order to prove the existence, we have to show that $\mathcal{S}_{1}$ maps the ball into itself, provided the constant is large enough, and that it is Lipschitz continuous in $V$ with Lipschitz constant of order $\varepsilon$. The Lipschitz dependence on the data is proved exploiting the Lipschitz character of $\mathrm{N}_{\varepsilon, \phi}, Q_{\varepsilon, \phi}$ and $\Gamma_{\varepsilon, \phi}$ (see (48), (50) and Remark 8) with respect to $U$ and $\phi$. More precisely, we use the fact that

$$
\begin{aligned}
&\left\|\mathrm{N}_{\varepsilon, \phi}\left(U_{1}\right)-\mathrm{N}_{\varepsilon, \phi}\left(U_{2}\right)\right\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{3}\right)} \leq c \mathrm{e}^{-\delta / 8 \varepsilon}\left\|U_{2}-U_{1}\right\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \\
&\left\|\mathrm{P}_{\varepsilon, \phi}\left(V_{\varepsilon, \phi, U_{1}}\right)-\mathrm{P}_{\varepsilon, \phi}\left(V_{\varepsilon, \phi, U_{2}}\right)\right\|_{C_{\delta}^{0, \alpha}\left(\mathbb{R}^{3}\right)} \leq c \mathrm{e}^{-\delta / 8 \varepsilon}\left\|V_{\varepsilon, \phi, U_{1}}-V_{\varepsilon, \phi, U_{2}}\right\|_{C_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \\
&\left\|\Gamma_{\varepsilon, \phi_{1}}-\Gamma_{\varepsilon, \phi_{2}}\right\|_{C_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leq c \mathrm{e}^{-\delta / 8 \varepsilon}\left\|\phi_{1}-\phi_{2}\right\|_{C^{4, \alpha}(\mathbb{R})}
\end{aligned}
$$

### 5.2. Proof of Proposition 10

The aim of this section is to solve equation (58). Recall that we defined (see (53))

$$
\mathcal{L}:=-\left(\partial_{s}^{2}+\partial_{t}^{2}\right)+W^{\prime \prime}\left(v_{0}(t)\right)
$$

We first consider the linear problem

$$
\begin{cases}\mathcal{L}^{2} U=f & \text { in } \mathbb{R}^{2}  \tag{76}\\ \int_{\mathbb{R}} U(s, t) v_{0}^{\prime}(t) \mathrm{d} t=0 & \forall s \in \mathbb{R}\end{cases}
$$

in order to produce a right inverse of $\mathcal{L}^{2}$ (see Section 5.1). Then we apply this right inverse to define a contraction on a suitable small ball that will give us the solution through a fixed point argument.

We recall that we endowed the spaces $D_{T, \delta}^{n, \alpha}\left(\mathbb{R}^{2}\right)$ with the weighted norms

$$
\|U\|_{D_{\delta}^{n, \alpha}\left(\mathbb{R}^{2}\right)}:=\left\|U \psi_{\delta}\right\|_{C^{n, \alpha}\left(\mathbb{R}^{2}\right)}
$$

where $\psi_{\delta}$ is defined in (44).
5.2.1. The Linear Problem As in Section 4, we first consider the second order problem

$$
\begin{cases}\mathcal{L} U=f & \text { in } \mathbb{R}^{2}  \tag{77}\\ \int_{\mathbb{R}} U(y, t) v_{0}^{\prime}(t) \mathrm{d} t=0 & \forall s \in \mathbb{R}\end{cases}
$$

In order to get an estimate in weighted norm, we need an a priori estimate, which we will state in the next Lemma. This result is similar to Lemma 6.2 in [18], but here the situation is simpler since we just have exponential weights on the $t$-variable, while in [18] there is also a weight in the limit manifold, which is non-periodic.

Lemma 16. (A priori estimate). Let $0<\delta<\sqrt{2}, f \in D_{T, \delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)$ and $U \in$ $D^{2, \alpha}\left(\mathbb{R}^{2}\right)$ be a solution to

$$
\begin{aligned}
& \mathcal{L} U=f \\
& \int_{\mathbb{R}} U(s, t) v_{0}^{\prime}(t) \mathrm{d} t=0, \quad \forall s \in \mathbb{R},
\end{aligned}
$$

satisfying

$$
U\left(x_{1}, x_{2}\right)=-U\left(-x_{1}, x_{2}\right)=-U\left(x_{1}+\frac{T}{2},-x_{2}\right), \quad \forall x=\left(x_{1}, x_{2}\right) \in \mathbb{R}^{2}
$$

and such that $U \psi_{\delta} \in L^{\infty}\left(\mathbb{R}^{2}\right)$. Then $U \in D_{T, \delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)$ and

$$
\begin{equation*}
\|U\|_{D_{\delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c\|f\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \tag{78}
\end{equation*}
$$

for some constant $c>0$ independent of $\varepsilon$.
Proof. As above, we set $\tilde{U}:=U \psi_{\delta}$ and $\tilde{f}:=f \psi_{\delta}$ (we recall that $\psi_{\delta}$ is a function of the $t$-variable). Since $\tilde{U}$ fulfils the equation

$$
-\Delta \tilde{U}+2 \psi_{-\delta} \partial_{t} \psi_{\delta} \partial_{t} \tilde{U}+\left(W^{\prime \prime}\left(v_{0}(t)\right)-2\left(\psi_{-\delta} \partial_{t} \psi_{\delta}\right)^{2}+\psi_{-\delta} \partial_{t t} \psi_{\delta}\right) \tilde{U}=\tilde{f}
$$

where $\psi_{-\delta}:=\psi_{\delta}^{-1}$, then by elliptic estimates it is enough to show that

$$
\left\|U \psi_{\delta}\right\|_{L^{\infty}\left(\mathbb{R}^{2}\right)} \leqq c\left\|f \psi_{\delta}\right\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}
$$

We argue by contradiction, that is we suppose that there exists a sequence $\varepsilon_{n} \rightarrow$ $0, T_{N}:=\varepsilon_{n}^{-1}, f_{n} \in D_{T_{n}, \delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)$ such that $\left\|f_{n}\right\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \rightarrow 0$ and a sequence $U_{n} \in$ $D_{T_{n}}^{2, \alpha}\left(\mathbb{R}^{2}\right)$ of solutions to

$$
\begin{align*}
\mathcal{L} U_{n} & =f_{n} \quad \text { in } \mathbb{R}^{2} ; \\
\int_{\mathbb{R}} U_{n}(s, t) v_{0}^{\prime}(t) \mathrm{d} t & =0 \forall s \in \mathbb{R}, \tag{79}
\end{align*}
$$

such that $\left\|U_{n} \psi_{\delta}\right\|_{L^{\infty}\left(\mathbb{R}^{2}\right)}=1$. In particular, there exists $\left(s_{n}, t_{n}\right) \in \mathbb{R}^{2}$ such that $\left|U_{n}\left(s_{n}, t_{n}\right)\right| \psi_{\delta}\left(t_{n}\right) \rightarrow 1$. We distinguish among three cases.
(i) First we assume that $\left|s_{n}\right|+\left|t_{n}\right|$ is bounded. By the uniform bound on the norms, up to a subsequence, $U_{n}$ converges in the $C_{l o c}^{2}\left(\mathbb{R}^{2}\right)$ sense to a bounded $C^{2}\left(\mathbb{R}^{2}\right)$-solution $U_{\infty}$ to

$$
\begin{equation*}
-\Delta U_{\infty}+W^{\prime \prime}\left(v_{0}(t)\right) U_{\infty}=0 \quad \text { in } \mathbb{R}^{2} \tag{80}
\end{equation*}
$$

Hence, by Lemma 6, 1 in [18], $U_{\infty}=\lambda v_{0}^{\prime}(t)$ for some $\lambda \in \mathbb{R}$. Moreover, by (79),

$$
\begin{equation*}
0=\int_{\mathbb{R}} U_{n}(s, t) v_{0}^{\prime}(t) \mathrm{d} t \rightarrow \int_{\mathbb{R}} U_{\infty}(s, t) v_{0}^{\prime}(t) \mathrm{d} t \quad \forall s \in \mathbb{R} \tag{81}
\end{equation*}
$$

thus $U_{\infty} \equiv 0$. However, up to a subsequence, $s_{n} \rightarrow s_{\infty}$ and $t_{n} \rightarrow t_{\infty}$, hence $\left|U_{n}\left(s_{n}, t_{n}\right)\right| \psi_{\delta}\left(t_{n}\right) \rightarrow\left|U_{\infty}\left(s_{\infty}, t_{\infty}\right)\right| \psi_{\delta}\left(t_{\infty}\right)=1$, a contradiction.
(ii) Now we assume that $t_{n}$ is unbounded. We set

$$
\tilde{U}_{n}(s, t):=U_{n}\left(s+s_{n}, t+t_{n}\right) \psi_{\delta}\left(t+t_{n}\right)
$$

As above, exploiting the equation satisfied by $\tilde{U}_{n}$, the uniform bound of the $L^{\infty}$ norms of $U_{n} \psi_{\delta}$ and elliptic estimates, we have that, up to a subsequence, $\tilde{U}_{n}$ converges in the $C_{\text {loc }}^{2}\left(\mathbb{R}^{2}\right)$ sense to a bounded solution $\tilde{U}_{\infty}$ to

$$
-\Delta \tilde{U}_{\infty}+2 \delta \partial_{t} \tilde{U}_{\infty}+\left(W^{\prime \prime}(1)-\delta^{2}\right) \tilde{U}_{\infty}=0
$$

By construction,

$$
\left|\tilde{U}_{n}(0,0)\right|=\left|U_{n}\left(s_{n}, t_{n}\right)\right| \psi_{\delta}\left(t_{n}\right) \rightarrow 1
$$

and $\left|\tilde{U}_{n}(s, t)\right| \leqq 1$ for any $(s, t) \in \mathbb{R}^{2}$, thus $\left|\tilde{U}_{\infty}(0,0)\right|=1=\sup _{(s, t) \in \mathbb{R}^{2}}\left|\tilde{U}_{\infty}\right|$. If, for instance, $\tilde{U}_{\infty}(0,0)=1$, then it is a maximum, hence

$$
\begin{aligned}
& \left(W^{\prime \prime}(1)-\delta^{2}\right)=\left(W^{\prime \prime}(1)-\delta^{2}\right) \tilde{U}_{\infty}(0,0) \\
& \quad \leq-\Delta \tilde{U}_{\infty}+2 \delta \partial_{t} \tilde{U}_{\infty}+\left(W^{\prime \prime}(1)-\delta^{2}\right) \tilde{U}_{\infty}=0
\end{aligned}
$$

a contradiction. With a similar argument, we can also exclude the case $\tilde{U}_{\infty}(0,0)=$ -1 .
(iii) It remains to rule out the case where $t_{n}$ is bounded and $s_{n}$ is unbounded. We define

$$
\tilde{U}_{n}(s, t):=U_{n}\left(s+s_{n}, t\right)
$$

As above, we have convergence, up to a subsequence to a bounded $C^{2}$ solution to (80). Since (81) is still true, once again we conclude that $\tilde{U}_{\infty} \equiv 0$. Nevertheless, extracting a subsequence $t_{n} \rightarrow t_{\infty}$ if necessary, we have

$$
\left|\tilde{U}_{n}\left(0, t_{n}\right)\right|=\left|U_{n}\left(s_{n}, t_{n}\right)\right| \rightarrow \psi_{-\delta}\left(t_{\infty}\right)
$$

hence $\left|\tilde{U}_{\infty}\left(0, t_{\infty}\right)\right|=\psi_{-\delta}\left(t_{\infty}\right)>0$, a contradiction.

Lemma 17. Let $0<\delta<\sqrt{2}$, and let $f \in D_{T, \delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)$ satisfy

$$
\begin{equation*}
\int_{\mathbb{R}} f(s, t) v_{0}^{\prime}(t) \mathrm{d} t=0 \quad \forall s \in \mathbb{R} \tag{82}
\end{equation*}
$$

Then, for $\varepsilon>0$ small enough, there exists a unique solution $U=\tilde{G}_{\varepsilon}(f) \in$ $D_{T, \delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)$ to (77) such that

$$
\|U\|_{D_{\delta}^{2, \alpha}\left(\mathbb{R}^{2}\right)} \leqq C\|f\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}
$$

for some constant $C>0$ independent of $\varepsilon$ and $\phi$.
Proof. Exploiting the periodicity, first we look for a weak solution $Z \in H^{1}(S)$ to the problem

$$
\begin{cases}-\Delta Z+W^{\prime \prime}\left(v_{0}(t)\right) Z=f & \text { in } S \\ \partial_{s} Z(-\pi / \varepsilon, t)=\partial_{s} Z(\pi / \varepsilon, t)=0 & \forall t \in \mathbb{R} \\ \int_{\mathbb{R}} Z(s, t) v_{0}^{\prime}(t) \mathrm{d} t=0 & \forall s \in(-T, T)\end{cases}
$$

where $S:=(-T, T) \times \mathbb{R}$, then we extend it to the whole $\mathbb{R}^{2}$. In other words, we look for a function $Z \in H^{1}(S)$ satisfying

$$
\begin{aligned}
\int_{S}\langle\nabla Z, \nabla v\rangle+\int_{S} W^{\prime \prime}\left(v_{0}(t)\right) Z v & =\int_{S} f v \forall v \in H^{1}(S) \\
\int_{\mathbb{R}} Z(s, t) v_{0}^{\prime}(t) \mathrm{d} t & =0 \quad \text { almost everywhere } s \in(-T, T)
\end{aligned}
$$

Since

$$
\int_{\mathbb{R}}\left(v^{\prime}\right)^{2}+W^{\prime \prime}\left(v_{0}\right) v^{2} \mathrm{~d} t \geqq c\|v\|_{H^{1}(\mathbb{R})}^{2}
$$

for any $v \in H^{1}(\mathbb{R})$ such that

$$
\int_{\mathbb{R}} v v_{0}^{\prime} \mathrm{d} t=0
$$

the symmetric bilinear form defined by

$$
b(Z, v):=\int_{S}\langle\nabla Z, \nabla v\rangle+\int_{S} W^{\prime \prime}\left(v_{0}(t)\right) Z v
$$

is coercive on the closed subspace
$X:=\left\{Z \in H^{1}(S): \int_{\mathbb{R}} Z(s, t) v_{0}^{\prime}(t) \mathrm{d} t=0 \quad\right.$ almost everywhere $\left.s \in(-T, T)\right\}$.
Therefore, by the Lax-Milgram theorem, there exists a unique $Z \in X$ such that

$$
\begin{equation*}
b(Z, v)=\int_{S} f v \quad \forall v \in X \tag{83}
\end{equation*}
$$

In order to show that $Z$ is actually a weak solution, we need to prove that (83) is true for any $v \in H^{1}(S)$. In order to do so, we decompose an arbitrary $v \in H^{1}(S)$ as

$$
v(s, t)=\tilde{v}(s, t)+c(s) v_{0}^{\prime}(t)
$$

where $a(s):=\int_{\mathbb{R}} v v_{0}^{\prime} \mathrm{d} t / \int_{\mathbb{R}}\left(v_{0}^{\prime}\right)^{2} \mathrm{~d} t$ is chosen in such a way that $\tilde{v} \in X$. We observe that, since

$$
\int_{S} f v_{0}^{\prime} \mathrm{d} t=0, \quad \forall s \in(-T, T)
$$

we have

$$
\int_{-T}^{T} c(s) \mathrm{d} s \int_{\mathbb{R}} f(s, t) v_{0}^{\prime}(t) \mathrm{d} t=0
$$

Moreover, an integration by parts and Fubini-Tonelli's Theorem show that

$$
b\left(Z, c v_{0}^{\prime}\right)=\int_{-T}^{T} \partial_{s} c(s) \partial_{s} \int_{\mathbb{R}} Z v_{0}^{\prime} \mathrm{d} t+\int_{-T}^{T} c(s) \int_{\mathbb{R}} Z L_{\star} v_{0}^{\prime} \mathrm{d} t=0
$$

In conclusion,

$$
b(Z, v)=b(Z, \tilde{v})+b\left(Z, c v_{0}^{\prime}\right)=\int_{S} f \tilde{v} \mathrm{~d} t=\int_{S} f v
$$

In order to prove symmetry and to extend $Z$ to an entire solution $U \in C^{2, \alpha}\left(\mathbb{R}^{2}\right)$, see Step (ii) of the proof of Lemma 14. Arguing as in Step (iii) of that proof, it is possible to show that $U \in L^{\infty}\left(\mathbb{R}^{2}\right)$. In order to show that $U \psi_{\delta} \in L^{\infty}\left(\mathbb{R}^{2}\right)$, we use the function $\lambda \mathrm{e}^{-\delta|t|}+\sigma \mathrm{e}^{\delta|t|}$ as a barrier, for suitable constants $\lambda$ and $\tau$. Here there is a slight difference with respect to the proof of Lemma 14, due the fact that the potential is not uniformly positive. This is actually not so relevant, since $W^{\prime \prime}\left(v_{0}(t)\right)$ is close to $W^{\prime \prime}(1)=2$ for $|t|$ large enough.

Now we can solve the fourth-order problem (76) by applying iteratively Lemma 17.

Lemma 18. Let $0<\delta<\sqrt{2}$ and let $f \in D_{T, \delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)$ satisfy (82). Then there exists a unique solution $U=G_{\varepsilon}(f) \in D_{T, \delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)$ to (76) such that

$$
\|U\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leqq C| | f \|_{D_{T, \delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)}
$$

for some constant $C>0$ independent of $\varepsilon$.
5.2.2. Proof of Proposition 10 Completed The proof is based on a fixed point argument. In fact, we have to find a fixed point of the map

$$
\begin{equation*}
\mathcal{S}_{2}(U):=G_{\varepsilon}\left\{-\chi_{4} F\left(\tilde{v}_{\varepsilon, \phi}\right)-\mathrm{T}\left(U, V_{\varepsilon, \phi, U}, \phi\right)+p_{\phi}(y) v_{0}^{\prime}(t)\right\} \tag{84}
\end{equation*}
$$

on a suitable small metric ball of the form
$\Lambda_{2}:=\left\{U \in D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right): \int_{\mathbb{R}} U(s, t) v_{0}^{\prime}(t) \mathrm{d} t=0 \quad \forall s \in \mathbb{R},\|U\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leqq C_{2} \varepsilon^{5}\right\}$,
provided that $C_{2}>0$ is large enough. Once again, we will prove that $\mathcal{S}_{2}$ is a contraction $\Lambda_{2}$. First we observe that, by definition of $p_{\phi}$, the quantity inside brackets in (84) is orthogonal to $v_{0}^{\prime}(t)$ for any $s \in \mathbb{R}$, thus we can actually apply the operator $G_{\varepsilon}$. Moreover, if $U$ respects the symmetries of the curve, then also the right-hand side does, hence applying $G_{\varepsilon}$ we get once again something that respects these symmetries.

In order to prove that $\mathcal{S}_{2}$ is a contraction, we note that

$$
\left\|F\left(\tilde{v}_{\varepsilon, \phi}\right)\right\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq \tilde{c} \varepsilon^{5}
$$

(see (47)), and a similar estimate is true for $p_{\phi}(s) v_{0}^{\prime}(t)$. The term $\mathrm{T}\left(U, V_{\varepsilon, \phi, U}, \phi\right)$ defined in (56) is smaller. For instance, using (48) and the fact that $V$ is exponentially small, one has that

$$
\left\|\chi_{1} Q_{\varepsilon, \phi}(U+V)\right\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c \varepsilon^{10}
$$

Similarly, we can see that $\left\|\mathbf{M}_{\varepsilon, \phi}(V)\right\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c \mathrm{e}^{-\delta / 4 \varepsilon}$. In addition, since all the coefficients of $\mathrm{R}_{\varepsilon, \phi}$ are at least of order $\varepsilon$, we get that

$$
\left\|\mathrm{R}_{\varepsilon, \phi}(U)\right\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c \varepsilon\|U\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c \varepsilon^{6}
$$

For the definitions of $\mathrm{M}_{\varepsilon, \phi}, R_{\varepsilon, \phi}$ and $Q_{\varepsilon, \phi}$, see (49), (52) and (48).
As regards the Lipschitz dependence on $U$, we observe that

$$
\left\|\chi_{1}\left(Q_{\varepsilon, \phi}\left(U_{1}+V\right)-Q_{\varepsilon, \phi}\left(U_{2}+V\right)\right)\right\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c \varepsilon^{5}\left\|U_{1}-U_{2}\right\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)}
$$

and

$$
\left\|\mathrm{R}_{\varepsilon, \phi}\left(U_{1}\right)-\mathrm{R}_{\varepsilon, \phi}\left(U_{2}\right)\right\|_{D_{\delta}^{0, \alpha}\left(\mathbb{R}^{2}\right)} \leqq c \varepsilon\left\|U_{1}-U_{2}\right\|_{D_{\delta}^{4, \alpha}\left(\mathbb{R}^{2}\right)}
$$

It follows from the Lipschitz character of the potential $W$ that the solution $U$ depends on $\phi$ in a Lipschitz way.

## 6. Appendix

In this section we provide a full proof of our claims from Proposition 4, and in particular of (16). To this end, we consider the solutions $\Phi(z)=\sum_{m=0}^{\infty} \mu_{m} z^{2 m+1}$ of the ODE (14). A coefficient comparison yields that $\mu_{2}, \mu_{3}, \ldots$ are explicitly given in terms of $\mu_{0}, \mu_{1}$ via the formulas

$$
\begin{align*}
\mu_{2 m} & =-\frac{3 \pi \sqrt{2}}{32 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \frac{\Gamma\left(m-\frac{1}{4}\right)}{4^{m} \Gamma\left(m+\frac{5}{4}\right)}-\frac{\mu_{0}}{\sqrt{\pi}} \cdot \frac{\Gamma\left(m+\frac{1}{2}\right)}{4^{m} \Gamma(m+1)(4 m-1)} \\
\mu_{2 m+1} & =\frac{3 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2}}{8 \pi} \cdot \frac{\Gamma\left(m+\frac{1}{4}\right)}{4^{m} \Gamma\left(m+\frac{7}{4}\right)} \mu_{1} \tag{85}
\end{align*}
$$

By the asymptotics of the Gamma function, see example [21], the convergence radius of this series is $\sqrt{2}$. Following the reasoning from Proposition 4 let us now derive two linear equations for $\mu_{0}, \mu_{1}$ so that any corresponding solution $\bar{\phi}=\Phi \circ k$ satisfies $\bar{\phi}^{\prime}(s), \bar{\phi}^{\prime \prime \prime}(s) \rightarrow 0$ as $|k(s)| \rightarrow \sqrt{2}$, that is as $|s| \rightarrow \bar{T} / 4$. It will turn out that the solution of this system is unique and given by $\mu_{0}=0, \mu_{1}=\frac{\pi^{2}}{8 \Gamma\left(\frac{3}{4}\right)^{4}}$, in accordance with (16).
Proof of (16). To this end we first calculate the derivatives of $\bar{\phi}$. For all $m \in \mathbb{N}_{0}$ we set

$$
\begin{aligned}
a_{m} & :=(2 m+1) \mu_{m} \\
b_{m} & :=-\frac{1}{2}(2 m+3)(2 m+1)(m+1) \mu_{m}+2(2 m+5)(m+2)(2 m+3) \mu_{m+2}
\end{aligned}
$$

Then, for all $s \in(-\bar{T} / 4, \bar{T} / 4)$ we have $|k(s)|<\sqrt{2}$ and thus we obtain from (6) and (9)

$$
\begin{aligned}
\bar{\phi}^{\prime}(s) & =\sum_{m=0}^{\infty} a_{m} k^{\prime}(s) k(s)^{2 m}, \\
\bar{\phi}^{\prime \prime}(s) & =\sum_{m=1}^{\infty} a_{m}\left(k^{\prime \prime}(s) k(s)^{2 m}+2 m k^{\prime}(s)^{2} k(s)^{2 m-1}\right)+a_{0} k^{\prime \prime}(s) \\
& =\sum_{m=1}^{\infty} a_{m}\left(-\frac{1}{2} k(s)^{2 m+3}+2 m k(s)^{2 m-1}\left(1-\frac{1}{4} k(s)^{4}\right)\right)-\frac{a_{0}}{2} k(s)^{3} \\
& =\sum_{m=0}^{\infty}\left(-\frac{1}{2}(m+1) a_{m}+2(m+2) a_{m+2}\right) k(s)^{2 m+3}+2 a_{1} k(s) \\
\bar{\phi}^{\prime \prime \prime}(s) & =\sum_{m=0}^{\infty} b_{m} k^{\prime}(s) k(s)^{2 m+2}+2 a_{1} k^{\prime}(s) .
\end{aligned}
$$

For the analysis of convergence, we rewrite $\bar{\phi}^{\prime}, \bar{\phi}^{\prime \prime \prime}$ as follows:

$$
\bar{\phi}^{\prime}(s)=k^{\prime}(s) \cdot \sum_{m=0}^{\infty}\left(a_{2 m}+2 a_{2 m+1}+a_{2 m+1}\left(k(s)^{2}-2\right)\right) k(s)^{4 m}
$$

$\bar{\phi}^{\prime \prime \prime}(s)=k^{\prime}(s) k(s)^{2} \cdot \sum_{m=0}^{\infty}\left(b_{2 m}+2 b_{2 m+1}+b_{2 m+1}\left(k(s)^{2}-2\right)\right) k(s)^{4 m}+2 a_{1} k^{\prime}(s)$.
Therefore we have to investigate the behaviour of the terms $a_{2 m}+2 a_{2 m+1}, b_{2 m}+$ $2 b_{2 m+1}, a_{2 m+1}, b_{2 m+1}$ as $m \rightarrow \infty$. To this end we use the known asymptotics (see p. 1 in [21])

$$
\begin{equation*}
\frac{\Gamma(z+\alpha)}{\Gamma(z+\beta)}=z^{\alpha-\beta} \cdot\left(1+\frac{(\alpha-\beta)(\alpha+\beta-1)}{2 z}+O\left(z^{-2}\right)\right) \text { as } z \rightarrow \infty \tag{86}
\end{equation*}
$$

for any fixed $\alpha, \beta \in \mathbb{R}$.
We start with analysing the behaviour of $\bar{\phi}^{\prime}(s)$ as $|s| \rightarrow \bar{T} / 4$. We have

$$
\begin{aligned}
a_{2 m}+2 a_{2 m+1}= & (4 m+1) \mu_{2 m}+2(4 m+3) \mu_{2 m+1} \\
= & (4 m+1) \cdot\left(-\frac{3 \pi \sqrt{2}}{32 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \frac{\Gamma\left(m-\frac{1}{4}\right)}{4^{m} \Gamma\left(m+\frac{5}{4}\right)}\right. \\
& \left.-\frac{\mu_{0}}{\sqrt{\pi}} \cdot \frac{\Gamma\left(m+\frac{1}{2}\right)}{4^{m} \Gamma(m+1)(4 m-1)}\right) \\
& +2(4 m+3) \cdot \frac{3 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2} \mu_{1}}{8 \pi} \cdot \frac{\Gamma\left(m+\frac{1}{4}\right)}{4^{m} \Gamma\left(m+\frac{7}{4}\right)} \\
= & \frac{1}{4^{m}} \cdot\left(-\frac{3 \pi \sqrt{2}}{8 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \frac{\Gamma\left(m-\frac{1}{4}\right)}{\Gamma\left(m+\frac{1}{4}\right)}-\frac{\mu_{0}}{\sqrt{\pi}} \cdot \frac{\Gamma\left(m+\frac{1}{2}\right)(4 m+1)}{\Gamma(m+1)(4 m-1)}\right. \\
& \left.+\frac{3 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2} \mu_{1}}{\pi} \cdot \frac{\Gamma\left(m+\frac{1}{4}\right)}{\Gamma\left(m+\frac{3}{4}\right)}\right) \\
= & \frac{1}{4^{m} m^{1 / 2}} \cdot\left(-\frac{3 \pi \sqrt{2}}{8 \Gamma\left(\frac{3}{4}\right)^{2}}-\frac{1}{\sqrt{\pi}} \cdot \mu_{0}+\frac{3 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2}}{\pi} \cdot \mu_{1}+O\left(m^{-1}\right)\right) \\
a_{2 m+1}= & \frac{3 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2} \mu_{1}}{2 \pi} \cdot \frac{\Gamma\left(m+\frac{1}{4}\right)}{4^{m} \Gamma\left(m+\frac{3}{4}\right)}=\frac{1}{4^{m}} \cdot O\left(m^{-1 / 2}\right) .
\end{aligned}
$$

Therefore we must require

$$
\begin{equation*}
-\frac{3 \pi \sqrt{2}}{8 \Gamma\left(\frac{3}{4}\right)^{2}}-\frac{1}{\sqrt{\pi}} \cdot \mu_{0}+\frac{3 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2}}{\pi} \cdot \mu_{1}=0 \tag{87}
\end{equation*}
$$

Once this equation is satisfied we have for some positive $C, C^{\prime}$ as $|s| \rightarrow \bar{T} / 4$

$$
\begin{aligned}
\left|\bar{\phi}^{\prime}(s)\right| & \leqq\left|k^{\prime}(s)\right| \cdot \sum_{m=0}^{\infty}\left(4^{m}\left|a_{2 m}+2 a_{2 m+1}\right|+4^{m}\left|a_{2 m+1}\right|\left|k(s)^{2}-2\right|\right)\left(\frac{k(s)^{4}}{4}\right)^{m} \\
& \leqq c\left|k^{\prime}(s)\right| \cdot\left(\sum_{m=0}^{\infty} m^{-3 / 2}+\sum_{m=0}^{\infty}\left|k(s)^{2}-2\right|\left(\frac{k(s)^{4}}{4}\right)^{m}\right) \\
& \leqq c^{\prime}\left|k^{\prime}(s)\right| \cdot\left(1+\left|k(s)^{2}-2\right| \cdot \frac{1}{1-\frac{k(s)^{4}}{4}}\right)=o(1)
\end{aligned}
$$

so the desired asymptotics for $\bar{\phi}^{\prime}$ holds. Hence, we have shown that any couple $\mu_{0}, \mu_{1}$ satisfying (87) yields the convergence of $\bar{\phi}^{\prime}(s)$ as $s \rightarrow \bar{T} / 4$.

Now we turn to the third-order derivatives. We have

$$
\begin{aligned}
b_{2 m}= & -\frac{1}{2}(4 m+3)(4 m+1)(2 m+1) \mu_{2 m}+2(4 m+5)(2 m+2)(4 m+3) \mu_{2 m+2} \\
= & \frac{3 \pi \sqrt{2}}{64 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \frac{(4 m+3)(4 m+1)(2 m+1) \Gamma\left(m-\frac{1}{4}\right)}{4^{m} \Gamma\left(m+\frac{5}{4}\right)} \\
& -\frac{3 \pi \sqrt{2}}{16 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \frac{(4 m+5)(2 m+2)(4 m+3) \Gamma\left(m+\frac{3}{4}\right)}{4^{m+1} \Gamma\left(m+\frac{9}{4}\right)} \\
& +\frac{\mu_{0}}{2 \sqrt{\pi}} \cdot \frac{(4 m+3)(4 m+1)(2 m+1) \Gamma\left(m+\frac{1}{2}\right)}{4^{m} \Gamma(m+1)(4 m-1)} \\
& -\frac{2 \mu_{0}}{\sqrt{\pi}} \cdot \frac{(4 m+5)(2 m+2)(4 m+3) \Gamma\left(m+\frac{3}{2}\right)}{4^{m+1} \Gamma(m+2)(4 m+3)} \\
= & \frac{3 \pi \sqrt{2}}{16 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \frac{1}{4^{m}}\left(\frac{(4 m+3)(2 m+1) \Gamma\left(m-\frac{1}{4}\right)}{\Gamma\left(m+\frac{1}{4}\right)}\right. \\
& \left.-\frac{(4 m+3)(2 m+2) \Gamma\left(m+\frac{3}{4}\right)}{\Gamma\left(m+\frac{5}{4}\right)}\right) \\
& +\frac{\mu_{0}}{2 \sqrt{\pi}} \cdot \frac{1}{4^{m}}\left(\frac{(4 m+3)(4 m+1)(2 m+1) \Gamma\left(m+\frac{1}{2}\right)}{(4 m-1) \Gamma(m+1)}\right. \\
& \left.-\frac{(4 m+5)(2 m+2) \Gamma\left(m+\frac{3}{2}\right)}{\Gamma(m+2)}\right)
\end{aligned}
$$

as well as

$$
\begin{aligned}
b_{2 m+1}= & -\frac{1}{2}(4 m+5)(4 m+3)(2 m+2) \mu_{2 m+1} \\
& +2(4 m+7)(2 m+3)(4 m+5) \mu_{2 m+3} \\
= & -\frac{3 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2} \mu_{1}}{16 \pi} \cdot \frac{(4 m+5)(4 m+3)(2 m+2) \Gamma\left(m+\frac{1}{4}\right)}{4^{m} \Gamma\left(m+\frac{7}{4}\right)} \\
& +\frac{3 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2} \mu_{1}}{4 \pi} \cdot \frac{4(4 m+7)(2 m+3)(4 m+5) \Gamma\left(m+\frac{5}{4}\right)}{4^{m+1} \Gamma\left(m+\frac{11}{4}\right)} \\
= & \frac{3 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2} \mu_{1}}{4 \pi} \cdot \frac{1}{4^{m}}\left(-\frac{(4 m+5)(2 m+2) \Gamma\left(m+\frac{1}{4}\right)}{\Gamma\left(m+\frac{3}{4}\right)}\right. \\
& \left.+\frac{(2 m+3)(4 m+5) \Gamma\left(m+\frac{5}{4}\right)}{\Gamma\left(m+\frac{7}{4}\right)}\right) .
\end{aligned}
$$

Using the asymptotics of the Gamma function, from (86) we obtain

$$
\begin{aligned}
b_{2 m}+2 b_{2 m+1} & =\frac{1}{4^{m} m^{1 / 2}} \cdot\left(\frac{9 \pi \sqrt{2}}{16 \Gamma\left(\frac{3}{4}\right)^{2}}+\frac{2}{\sqrt{\pi}} \cdot \mu_{0}-\frac{9 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2}}{2 \pi} \cdot \mu_{1}+O\left(m^{-1}\right)\right), \\
b_{2 m+1} & =\frac{1}{4^{m}} \cdot O\left(m^{-1 / 2}\right)
\end{aligned}
$$

This leads us to require

$$
\begin{equation*}
\frac{9 \pi \sqrt{2}}{16 \Gamma\left(\frac{3}{4}\right)^{2}}+\frac{2}{\sqrt{\pi}} \cdot \mu_{0}-\frac{9 \sqrt{2} \Gamma\left(\frac{3}{4}\right)^{2}}{2 \pi} \cdot \mu_{1}=0 \tag{88}
\end{equation*}
$$

As before we obtain that every $\mu_{0}, \mu_{1}$ satisfying (88) makes sure that $\bar{\phi}^{\prime \prime \prime}(s)$ tends to zero as $s \rightarrow \bar{T} / 4$, that is as $|k(s)| \rightarrow \sqrt{2}$.

Collecting all the above reasoning, we find that $\bar{\phi}^{\prime}(s), \bar{\phi}^{\prime \prime \prime}(s) \rightarrow 0$ as $|k(s)| \rightarrow$ $\sqrt{2}$ provided $\mu_{0}, \mu_{1}$ solve (87), (88), which is a linear system with a unique solution $\mu_{0}=0, \mu_{1}=\frac{\pi^{2}}{8 \Gamma\left(\frac{3}{4}\right)^{4}}$. Plugging these values into the formula for $\bar{\phi}(s)=\Phi(k(s))$ we get
$\bar{\phi}(s)=\frac{3 \pi \sqrt{2}}{64 \Gamma\left(\frac{3}{4}\right)^{2}} \cdot \sum_{m=0}^{\infty}\left(-\frac{\Gamma\left(m+\frac{3}{4}\right)}{2 \cdot 4^{m} \Gamma\left(m+\frac{9}{4}\right)} k(s)^{4 m+5}+\frac{\Gamma\left(m+\frac{1}{4}\right)}{4^{m} \Gamma\left(m+\frac{7}{4}\right)} k(s)^{4 m+3}\right)$,
which is precisely the formula from Proposition 4.

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